## **Magnetic Fields in HED and ICF Systems**

**ICF** Lecture

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#### Limited Distribution

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### Magnetic fields are everywhere! Even ICF







#### Magnetic fields are everywhere! Even ICF





## **Outline**

#### FUNDAMENTALS

- Particle motion, drifts, magnetic mirroring ٠
- Waves in magnetized plasma, Faraday rotation ۰
- MHD single-fluid model: derivation ۰
- MHD effects •
  - Magnetic pressure
  - Reduced heat flow •
  - Frozen-in law
  - Biermann battery
  - Nernst, Righi-Leduc •

#### **OTHER DOMAINS not covered**

- Astrophysics, lab-astro
- Plasma propulsion

#### **EXPERIMENTS / APPLICATIONS**

- Diagnostics
- Imposed fields pulsed power, laser-driven coils
- Flux compression explosives, laser driven
- Fusion •
  - Magnetic vs. inertial fusion
  - Magneto-inertial fusion
  - Imposed fields in ICF
  - MagLIF
- B fields in hohlraums: self-generated, imposed





# Thanks for slides, conversations, insight, codes, data, supervision ...

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## **Motion in uniform B field: gyromotion**

- Lorentz force:  $\frac{d\vec{p}}{dt} = q\vec{v} \times \vec{B}$   $\vec{p} = \gamma m\vec{v}$   $\gamma \equiv (1 v^2/c^2)^{-1/2}$ Magnetic fields do no work:  $\vec{p} \cdot \frac{d\vec{p}}{dt} = 0 \rightarrow \gamma = \text{constant}$   $\vec{B} = B\hat{z}, \quad B > 0 \rightarrow \frac{d\vec{v}}{dt} = \omega_c \vec{v} \times \hat{z}$   $\vec{k} = 0$   $\vec{k} = 0$   $\vec{k} = 0$   $\vec{k} = 0$

$$\vec{B} = B\hat{z}, \qquad B > 0 \quad \rightarrow \quad \frac{d\vec{v}}{dt} = \omega_c \vec{v} \times \hat{z}$$

- Cyclotron frequency:  $\omega_c \equiv \frac{qB}{vm}$
- Initial condition:  $\vec{x}(t=0) = \vec{0}$ ,  $\vec{v}(t=0) = (v_{r0}, 0, v_z)$

$$\vec{x} = \frac{v_{x0}}{\omega_c} (\sin \omega_c t, \cos \omega_c t - 1, 0) + v_z t \hat{z}$$

Gyroradius:

$$R \equiv \left| \frac{v_{x0}}{\omega_c} \right| = \frac{(2mW)^{\frac{1}{2}}}{qB} \sin \theta \left( 1 + \frac{W}{2mc^2} \right)^{1/2}$$
  
relativity



 $\overrightarrow{B}||\hat{z}$ 

Helical orbit





#### Some numbers

	B [T]	Kinetic Energy	Cyclotron period [ps]	Gyroradius [um]	Example
Electron	10	1 keV	3.58	10.7	MIFEDS, MFE
Hydrogen	10	1 keV	6560	457	Same
Alphas (He4)	27,000	3.5 MeV	4.83 <~ hotspot size /	10.0 burn duration	Implosion: 30 T @ CR=30

#### **Electrons**

cyclotron period 
$$\tau_{ce} \equiv \frac{2\pi}{\omega_c} = 36 \text{ps} \frac{\gamma}{B[\text{T}]}$$
  
gyroradius  $R = 107 \mu \text{m} \frac{(W[\text{keV}])^{\frac{1}{2}}}{B[\text{T}]} \sin\theta \left(1 + \frac{W}{1 \text{ MeV}}\right)^{1/2}$ 



## Particle drifts in B field: $\vec{E} \times \vec{B}$ drift

y [V<sub>0</sub>/ω<sub>c</sub>]

$$\frac{d}{dt}(v_x, v_y) = \left(\frac{q}{m}E + \omega_c v_y, -\omega_c v_x\right)$$

E increases  $v_x \rightarrow v_y$  more negative  $\rightarrow$  reduces  $dv_x/dt$ 





Retional Nuclear Security Administration

## Particle drifts in B field: $\nabla \vec{B}$ , curvature, etc.

#### More on $\overrightarrow{E} \times \overrightarrow{B}$ drift

- Same for all particles independent of mass, charge, sign of charge
- Perpendicular to both  $\vec{B}$  and "driving force"  $\vec{E}$  in this case
  - Common aspect of magnetic effects, e.g. Nernst, Righi-Leduc effects
- Zoo of other drifts
  - Grad B
  - Curvature important for tokamaks
  - Polarization
  - Most are perpendicular to  $\vec{B}$  and a "driving force"

See Jackson or any plasma physics textbook, e.g. Krall and Trivelpiece

 Free online book: Prof. R. Fitzpatrick (UT Austin): farside.ph.utexas.edu/teaching/plasma/plasma.html



# Motion parallel to $\overrightarrow{B}$ : adiabatic invariance of magnetic moment

Action for periodic motion in q<sub>i</sub>:

$$J_i = \oint p_i \, dq_i$$
one cycle

 $q_i$  = canonical coordinate  $p_i$  = canonical momentum

In B field: canonical momentum  $\vec{p} = m\vec{v} + q\vec{A}$ 

Strictly periodic motion: J conserved

Nearly periodic motion: J approximately (but very well) conserved

Adiabatic invariance: parameters change slowly vs. period (and not resonant with it )

 $\vec{B}$  slowly varying in space: change over one gyro-orbit small:

Transverse motion  $\approx$  periodic  $\rightarrow$ 

$$J = \oint \vec{p}_{\perp} \cdot d\vec{l} = qB\pi R^2 \propto \frac{v_{\perp}^2}{B}$$

flux through gyro-orbit

$$\therefore$$
 Magnetic moment  $\mu \equiv \frac{v_{\perp}^2}{B}$  invariant

Non-relativistic



I. Hutchinson, MIT



## Motion parallel to $\overrightarrow{B}$ : magnetic mirror

Static, non-uniform  $\vec{B}$ ;  $\vec{E} = 0$ Kinetic energy conserved:  $v_{\perp}^2 + v_{||}^2 = v_{\perp 0}^2 + v_{||0}^2$ Magnetic moment adiabatic invariant:  $\frac{v_{\perp}^2}{B} = \frac{v_{\perp 0}^2}{B_0}$ Magnetic mirror: turning point:  $v_{||} = 0 \Rightarrow \frac{B_{mir}}{B_0} = 1 + \frac{v_{\perp 0}^2}{v_{\perp 0}^2}$ Loss cone in velocity: particles escape if no turning point  $\Rightarrow \theta < \arctan\left[\left(\frac{B_{max}}{B} - 1\right)^{1/2}\right]$   $\frac{F_{\perp}}{B_0}$  confined  $\frac{F_{\perp}}{B_0}$ "Magnetic mirror": MEE concept



"Magnetic mirror": MFE concept Loss-cone hard to overcome TMX: Tandem Mirror Experiment LLNL: 1979-1987 TMX (near superblock)





#### Waves in magnetized plasmas





### Waves in magnetized plasma: a zoo

- $\vec{E} = \vec{E}_{in} + \vec{E}_{ex}$  internal (plasma, "bound") + external (driven, "free") fields
- Linearize + Fourier transform { Maxwell's equations + plasma equations of motion }  $\rightarrow$
- Dispersion relation:  $\vec{D} \cdot \vec{E}_{in} = \overleftrightarrow{\chi} \cdot \vec{E}_{ex}$   $\overleftrightarrow{\chi}$  = susceptibility tensor

$$\vec{D} = \vec{\epsilon} + \vec{N}\vec{N} - N^2 \qquad \vec{N} \equiv c\vec{k}/\omega$$

- Normal modes:  $\vec{E}_{in} \neq 0$  when  $\vec{E}_{ex} = 0$ :  $\rightarrow$   $\vec{D} \cdot \vec{E}_{in} = 0 \rightarrow \det \vec{D} = 0$
- Cold plasma:

dielectric 
$$\vec{\epsilon} = 1 + \vec{\chi} = \begin{bmatrix} S & -iD & 0\\ iD & S & 0\\ 0 & 0 & P \end{bmatrix}$$
  
 $B = B\hat{z}$   
 $R = 1 - \sum_{s} \frac{\omega_{ps}^2}{\omega(\omega \pm \omega_{cs})}$   
 $S = \frac{1}{2}(R + L)$   
 $D = \frac{1}{2}(R - L)$   
 $P = 1 - \frac{\sum_{s} \omega_{ps}^2}{\omega^2}$ 

Classic text: T. H. Stix, Waves in Plasmas

Alfven waves, Bernstein waves, electrostatic ion cyclotron waves, O mode, X mode, ...



### Wave dispersion relation



Fixed ions, cold electrons,  $B=0 \rightarrow 3$  modes:

• R = L = S = P, D=0 •  $\vec{D} = \begin{bmatrix} P - N^2 \cos^2 \theta & 0 & N^2 \cos \theta \sin \theta \\ 0 & P - N^2 & 0 \\ N^2 \cos \theta \sin \theta & 0 & P - N^2 \sin^2 \theta \end{bmatrix}$   $\theta$  arbitrary, can set to 0

- det  $\vec{D} = (P N^2)^2 P = 0$  independent of  $\theta$
- P=0: Plasma oscillation, electrostatic  $(\vec{E} || \vec{k})$ :  $\omega = \omega_{pe}$
- P-N<sup>2</sup>=0: Electromagnetic (light,  $\vec{E} \perp \vec{k}$ ): 2 polarizations:  $\omega^2 = \omega_{pe}^2 + c^2 k^2$



## $\vec{k}||\vec{B}: \theta = 0$ principal modes

Fixed ions, cold electrons,  $B \neq 0 \rightarrow 4$  modes:

- Plasma oscillation  $(\vec{E} | | \vec{k})$ :  $\omega = \omega_{pe}$
- Electromagnetic  $(\vec{E} \perp \vec{k})$ : 3 modes:

$$\omega^3 \pm \omega_{ce}\omega^2 - (\omega_{pe}^2 + c^2k^2)\omega \mp \omega_{ce}c^2k^2 = 0$$

$$\label{eq:ce} \begin{split} \omega_{ce} &= 0.5 \omega_{pe} \\ \text{e.g. } n_{\text{e}} \text{=} 10^{20} \text{ cm}^{\text{-3}} \text{, } \text{B} \text{=} 1600 \text{ T} \end{split}$$







## $\vec{k} || \vec{B} || \hat{z}$ : polarization, Faraday rotation







## <sup>1</sup>Only two names. Declined knighthood, preferred to remain "plain Mr. Faraday to the end" nce Livermore National Laboratory

#### **Faraday rotation: First experimental evidence** that light and electromagnetism are related

Michael Faraday<sup>1</sup>, 1845

Faraday's diary, 1845:

"...when the contrary magnetic poles were on the same side, there was an effect produced on the polarized ray, and thus magnetic force and light were proved to have relation to each other."

"I have at last succeeded in illuminating a magnetic curve or line of force, and in magnetizing a ray of light."

Heavy glass (lead traces)





Magnetic: external waves (e.g. microwaves) used to drive current and heat plasma

Inertial: we use lasers! Laser-plasma interaction should include B fields

- B. Winjum (UCLA): PIC simulations
- E. Los (LLNL summer student): analytic instability growth / gain
- N. Fisch et al.: parametric amplifiers with magnetized waves



#### **MHD single-fluid model**





## MHD "single-fluid" model: equivalent to one ion, one electron fluid species

 $\rho = m_i n_i + m_e n_e \qquad \text{mass density}$   $Q = Z_i e n_i - e n_e \qquad \text{charge density}$   $\vec{v} = \frac{m_i n_i \vec{v}_i + m_e n_e \vec{v}_e}{\rho} \qquad \text{center - of - mass velocity}$   $\vec{J} = Z_i e n_i \vec{v}_i - e n_e \vec{v}_e \qquad \text{current density}$ 



#### Single-fluid equations: plasma + Maxwell

Plasma: 
$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0$$
$$\frac{\partial Q}{\partial t} + \nabla \cdot \vec{J} = 0$$
$$\rho \frac{D \vec{v}}{D t} = -\nabla p + Q \vec{E} + \vec{J} \times \vec{B}$$
$$\frac{1}{\epsilon_0 \omega_{pe}^2} \frac{D \vec{J}}{D t} = \vec{E} - \vec{E}_{eff}$$

+ energy equation

mass continuity

charge continuity

CM momentum

Ohm's law  $\approx e^{-}$  momentum

$$\omega_{pe}^2 = \frac{n_e e^2}{\epsilon_0 m_e}$$
 plasma frequency

Maxwell:



# Scaled variables: $X = X_0 \widetilde{X}$ , $X_0 =$ "unitful" scale, $\widetilde{X}$ =unitless

$$t = T\tilde{t} \qquad \vec{x} = L\tilde{\vec{x}} \qquad \vec{v} = V\tilde{\vec{v}} \qquad \text{Unitless scales}$$

$$n_e = N_e \widetilde{n_e} \qquad \rho = \frac{m_i N_e}{Z_i} \widetilde{\rho} \qquad Q = e N_e \alpha_Q \widetilde{Q} \qquad \vec{J} = e N_e V \alpha_J \tilde{\vec{J}}$$

$$\vec{E} = E_0 \tilde{\vec{E}} \qquad \vec{B} = B_0 \tilde{\vec{B}} \qquad \alpha_Q \ll 1: \text{ quasi-neutral}$$

#### **Balance terms:**

$$\begin{array}{lll} \text{Mass continuity:} & \frac{\partial \tilde{\rho}}{\partial \tilde{t}} + \frac{VT}{L} \tilde{\nabla} \cdot \left( \tilde{\rho} \tilde{\vec{v}} \right) = 0 & \rightarrow & V = \frac{L}{T} \\ \\ \text{Faraday:} & \frac{\partial \tilde{\vec{B}}}{\partial \tilde{t}} = \frac{TE_0}{LB_0} \tilde{\nabla} \times \tilde{\vec{E}} & \rightarrow & B_0 = \frac{T}{L} E_0 \end{array}$$

Homework: Turn the crank...





### Scaled single-fluid model: unitless, drop ~'s

Plasma: $\frac{\partial \rho}{\partial t} + \nabla \cdot (\alpha_Q \frac{\partial Q}{\partial t} + \alpha_J)$	$(\rho \vec{v}) = 0$ $\nabla \cdot \vec{J} = 0$	mass continuity charge continuity
M <sup>2</sup> $\mu$	$p\frac{D\vec{v}}{Dt} = -\nabla p + \beta_p^{-1}(\vec{J} \times \vec{B})$	$+ \beta_V^2 Q \vec{E}$ ) CM momentum
Maxwell:	$\vec{E} = \vec{E}_{eff} + \frac{1}{\left(\beta_V \omega_{pe} T\right)^2}$	$\frac{1}{n_e} \frac{D\vec{J}}{Dt}$ Ohm's law
$\nabla \cdot \vec{E} = Q$ $\nabla \cdot \vec{B} = 0$	Gauss	<u>6 Unitless parameters</u>
$\frac{\partial \vec{B}}{\partial t} = 0$ $\frac{\partial \vec{B}}{\partial t} = -\nabla \times \vec{E}$	Faraday	$\begin{array}{cc} \alpha_Q, & \alpha_J \\ \beta_V \equiv L/cT \end{array}$
$\nabla \times \vec{B} = \vec{J} + \beta_V^2 \frac{\partial E}{\partial t}$	Ampère – Maxwell	$\omega_{pe}T$ $M \equiv V/c_s  \text{Mach number}$
<b>MHD ordering:</b> $\beta_V = \frac{L}{cT} \ll 1$ drop respectively.	ed terms	$c_s^2 = \frac{Z_i p_0}{m_i N_e} \text{ sound speed}$ $\beta_p \equiv \frac{\mu_0 p_0}{B_0^2}  (1/2) \text{ plasma beta}$
$\omega_{pe} I \gg \frac{1}{\beta_V}$ drop blu	le term (e mertia)	
$\rightarrow \omega_{pe}T \gg \frac{CI}{L} \gg 1$		Light and Langmuir waves eliminated: Model does not support them





# Homework: what about quasi-neutrality and charge continuity?

- Turn the crank
- What is the relation between  $\alpha_Q$  and  $\alpha_I$ ?
- How does this relate to charge continuity?
  - What "loophole" allows the MHD model to satisfy it?





## Single-fluid MHD model: physical units

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0$$

$$\rho \frac{D \vec{v}}{Dt} = -\nabla p + \vec{J} \times \vec{B}$$
Magnetic pressure:  

$$\vec{J} \times \vec{B} = -\nabla \left(\frac{B^2}{2}\right) + \vec{B} \cdot \nabla \vec{B}$$
Maxwell:  

$$\nabla \cdot \vec{B} = 0$$

$$\partial \vec{B} / \partial t = -\nabla \times \vec{E}$$

$$\vec{J} = \mu_0^{-1} \nabla \times \vec{B}$$
Generalized  
Dhm's law:  

$$\vec{E} = \vec{E}_{eff} = -\vec{v} \times \vec{B} + \frac{1}{n_e e} \vec{J} \times \vec{B} - \frac{\nabla p_e}{n_e e} + \vec{\eta} \cdot \vec{J} - e^{-1} \vec{\beta} \cdot \nabla T_e$$
advection / Hall term Biermann resistivity thermal force\*  
induction term battery  
collisionless collisional

- E field from Ohm's law: function of current plasma conditions
- B field evolves in time by Faraday's law not function of current plasma conditions
- MHD models distinguished by terms in Ohm's law:
  - Ideal MHD: just advection
  - Resistive MHD: also resistivity
  - Hall MHD: also Hall term
- In ICF /HED, hard to say a priori which terms can be neglected for all time and space

\*thermal force collisional, but  $\overleftrightarrow{\beta}(B=0)$  independent of collisionality



### **Magnetic field generation: Biermann battery<sup>1</sup>**

Electron EOS:  $p_e = n_e T_e$  neglects Fermi degeneracy

Biermann battery: magnetic fields "spontaneously" (from B=0 state) develop due to nonparallel electron density and temperature gradients



<sup>1</sup>Ludwig Franz Benedict Biermann, Zeitschrift fur Naturforschung 5: 65 (1950)



### Magnetic field evolution: advection, diffusion



- Advection: B field advected with plasma flow, frozen-in law (2 slides below)
- Diffusion: breaks frozen-in law, dissipates magnetic energy
  - Magnetic Reynolds number: field advection vs. diffusion
  - Sets time to diffuse through good conductor, e.g. magnetizing a hohlraum with external coils
- $\nabla \eta$ : useful for schemes to self-guide electron beams for e.g. fast ignition:
  - A. Robinson et al.
- Imposed B fields also considered for fast ignition:
  - D. J. Strozzi, M. Tabak, et al., Phys. Plasmas 2012: compressed in implosion
  - Osaka group: laser-driven <~ 1 kT</li>



## MHD effects in hohlraums<sup>1</sup>

#### Slide courtesy Will Farmer



<sup>1</sup>W A Farmer, J M Koning, D J Strozzi, D E Hinkel, L F Berzak Hopkins, O S Jones, M D Rosen, Phys. Plasmas (2017)



## **Frozen-in law, or Alfven's theorem<sup>1</sup>**

Ideal MHD:  $\vec{E} = -\vec{v} \times \vec{B}$ , magnetic flux though surface moving with plasma conserved

$$\Phi \equiv \int_{S} \vec{B} \cdot d\vec{a} \qquad \text{magnetic flux through surface S}$$

$$\frac{\partial \Phi}{\partial t} = \int_{S} (\partial_{t}\vec{B} + (\nabla \cdot \vec{B})\vec{v}) \cdot d\vec{a} - \oint_{\partial S} \vec{v} \times \vec{B} \cdot d\vec{l} \qquad \text{Leibniz's theorem}$$

$$\frac{\partial \Phi}{\partial t} = \int_{S} \nabla \times (-\vec{E}) \cdot d\vec{a} - \oint_{\partial S} \vec{v} \times \vec{B} \cdot d\vec{l} \qquad \text{Faraday's law}$$

$$= \oint_{\partial S} -(\vec{E} + \vec{v} \times \vec{B}) \cdot d\vec{l} \qquad \text{Stokes' theorem} \qquad \text{Completely general}$$

$$= 0 \qquad \vec{E} = -\vec{v} \times \vec{B} \text{ on boundary}$$
Good conductor
$$\vec{V} \qquad \vec{B} \qquad \text{time} \qquad \text{State 2} \qquad B_{2} = \left(\frac{R_{1}}{R_{2}}\right)^{2} B_{1} \qquad \text{for all of the completely general}$$

$$= 1 \qquad \text{Hannes Olof Gösta Alfvén:}$$

$$= tc \qquad \text{"Existence of electromagnetic-hydromagnetic waves," Nature (1942)}$$



# Collisional tensors with B field: resistivity, thermal conductivity

**Resistivity**: 
$$\hat{\eta} = \begin{bmatrix} \eta_{\perp} & \eta_{\wedge} & 0 \\ -\eta_{\wedge} & \eta_{\perp} & 0 \\ 0 & 0 & \eta_{||} \end{bmatrix}$$
  $\vec{B} = b\hat{z}$   
 $\vec{\beta}$  and  $\vec{\kappa}$  same form – watch out for sign of ^ term!  
"Hall effect": Electric field develops perpendicular to applied B field and current  
 $\vec{J} = J\hat{y} \rightarrow \vec{E} = \vec{\eta} \cdot \vec{J} = (\eta_{\wedge}J, \eta_{\perp}J, 0) = \eta_{\perp}\vec{J} + \eta_{\wedge}\vec{J} \times \hat{b}$ 

Combines with collisionless JxB term

#### Heat flow reduced perpendicular to B

$$\vec{q} = \overleftarrow{\kappa} \cdot \nabla T_e$$

Hall parameter  $\chi \equiv \omega_{ce} \tau_{ei}$ Characterizes magnetization of transport coeffs.

 $\kappa(Z, \chi)$  found numerically, e.g. Epperlein and Haines, Phys. Fluids 1986

Don't forget Righi-Leduc! Heat flow along  $\vec{B} \times \nabla T_e$ 





#### **HED magnetic field experiments**





# Magnetic field measurements in laser-produced plasmas date to at least 1971

**Biermann fields** 



- 1971 J. A. Stamper et al., Phys. Rev. Lett.:
   0.1 T fields from B-dot probes
- 1975 J. A. Stamper, B. H. Ripin, Phys. Rev. Lett.: 100 T fields from Faraday rotation
- 1991 J. A. Stamper, Laser and Particle Beams: Review article



## Magnetic field diagnostics in HED

B-dot probes: workhorse in magnetic confinement devices

• Hard to use in HED / laser-plasma systems

#### Faraday rotation: line-integral of B\*n<sub>e</sub>

• Need path where n<sub>e</sub> sub-critical for probe wavelength

#### Proton radiography:

- 1-100 MeV protons feel E and B fields, and collisions with background
- Line-integral of Lorentz force
- Hard to distinguish E and B:
  - Multiple energies, geometric inversion
- Short-pulse laser protons: TNSA, other ion acceleration schemes
  - Borghesi et al., Plasma Phys. Control. Fusion 2001, Phys. Plasmas 2002
  - MacKinnon et al., Phys. Rev. Lett. 2006
- Fusion protons (mono-energetic): D-He3 exploding pusher
  - C. K. Li, R. Petrasso numerous paper

#### **Spectroscopy: Zeeman splitting** $\Delta E = \pm \frac{1}{2} \hbar \omega_{ce} = 0.058 \text{ eV/kT}$

- Measured imposed ~ 10 T field on Janus via optical Neon lines
- B. B. Pollock, D. H. Froula, P. F. Davis et al., Rev. Sci. Sci. 2006



#### **Proton radiography: D-He3 protons**

PRL 103, 085001 (2009) PHYSIC

PHYSICAL REVIEW LETTERS

week ending 21 AUGUST 2009

#### Lorentz Mapping of Magnetic Fields in Hot Dense Plasmas

R. D. Petrasso, C. K. Li, F. H. Seguin, J. R. Rygg,<sup>\*</sup> and J. A. Frenje Plasma Science and Fusion Center, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

> R. Betti, J. P. Knauer, and D. D. Meyerhofer Laboratory for Laser Energetics, University of Rochester, Rochester, New York 14623, USA

> P. A. Amendt, D. H. Froula, O. L. Landen, P. K. Patel, J. S. Ross, and R. P. J. Town Lawrence Livermore National Laboratory, Livermore, California 94551, USA (Received 12 March 2009; published 17 August 2009)

Unique detection of electromagnetic fields and identification of field type and strength as a function of position were used to determine the nature of self-generated fields in a novel experiment with laser-generated plasma bubbles on two sides of a plastic foil. Field-induced deflections of monoenergetic 15-MeV probe protons passing through the two bubbles, measured quantitatively with proton radiography, were combined with Lorentz mapping to provide separate measurements of magnetic and electric fields. The result was absolute identification and measurement of a toroidal magnetic field around each bubble and determination that any electric field component parallel to the foil was below measurement uncertainties.





Imposed fields: pulsed-power laser-driven coils





### Imposed B field: pulsed power: 10 – 40 T

- $B_0 = 12 \text{ T}$  at Janus (LLNL):
  - e- heat flow reduced to Braginskii values, nonlocality for B<sub>0</sub> = 0 quenched
  - B. B. Pollock, D. H. Froula, P. F. Davis et al., Rev. Sci. Sci. 2006
  - D. H. Froula, J S Ross, B B Pollock et al., Phys. Rev. Lett. 2007
- MIFEDS at Omega: 2007 present:
  - Magneto-Inertial Fusion Electrical Discharge System
  - 8-10 T, microsec. risetime, plans to upgrade
  - Cylindrical and spherical implosions, mini MagLIF
  - 2015: magnetized hohlraum: D. Montgomery et al.
- 2013 present: Z machine (Sandia) for MagLIF
  - Split Helmholtz coil: 10 T w/ diagnostic access, 30 T without
  - Very slow rise time vs. MIFEDS: several milliseconds, to diffuse through Be liner
- 2013: B<sub>0</sub> = 40 T at ELFIE laser (Ecole Polytechnique)
  - Albertazzi et al., Rev. Sci. Instrum. 2013
- 2014 present: U. Michigan coil, used on Janus, 12.5 T
  - Used for astrophysical magnetized jets, E. Kemp mid-Z foams
  - S. R. Klein, M. J.-E. Manuel, B. B. Pollock et al., Rev. Sci. Inst. (2014)
  - M.J.-E. Manuel, C.C. Kuranz, A.M. Rasmus, High. En. Dens. Phys. (2015)





## Imposed magnetic field: laser-driven coil

Laser-driven hot electrons carry current b/t 2 conductors



Goyon et al., 2017



1979 V V Korobkin, S L Motylev: Sov. Technical Phys. Lett.

First paper on basic concept

1986 H Daido et al., Phys. Rev. Lett. (Osaka)

• 60 T, 10.6 um CO<sub>2</sub> laser

2005 C Courtois et al., J. Applied Phys. (York)

- 7 T, Nd:glass laser (1.053 um), hot electron model
- 2013 S Fujioka et al., Scientific Reports (Osaka)
  - 1500 T (!!), 0.53 um or 1.06 um Nd:glass laser
  - Subsequent analysis suggests << field

2015 J J Santos et al., New J. Phys. (LULI pico 2000)

- 800 T
- B-dot probes, Faraday rotation, proton radiography

2017 C Goyon et al., Phys. Rev. E (Omega EP)

- 210 T, 0.351 um Nd:glass laser
- Time and space dynamics, lumped circuit model



Flux compression: explosive drive laser drive



#### **Flux compression with explosives**

#### Pioneers of magnetic flux compression

USSR VNIIEF - Sarov



Andrei Sakharov 1921-1989 USA Los Alamos



C. Max Fowler 1918-2006

J. L. Fowler, Los Alamos (1944, unpublished) ~ experiment, ~hundreds gauss.

Ya. P. Terletskii, Sov. Phys. JETP 32, 301 (1957) - theory, original report dated 1952.

C. M. Fowler, W. B. Garn and R. S. Caird, J. Appl. Phys. 31, 588 (1960) - explosive experiments, peak 15 MG.

A. D. Sakharov et al., Dokl. Akad. Nauk SSSR 165, 65 (1965); A. D. Sakharov, Usp. Fiz. Nauk 88, 725 (1966).

Suggested in 1951, explosive experiments started in 1952, peak 25 MG.

C.M. Fowler and L.L. Altgilbers, "Magnetic Flux Compression Generators: a Tutorial and Survey," Electromagnetic Phenomena **3**, No. 3 (11), 306 (2003).

Slide courtesy A. Velikovich



## Flux compression with explosives



#### Literature

1944: J. L. Fowler, unpublished – LANL

1960 LANL: C. M. Fowler, W. Garn, R. Caird, J. Applied Phys.: 1400 T

1965: A. D. Sakharov, Soviet Phys. Uspekhi

Accelerate particles to 1 TeV - nuclear explosive driven

1993: C. M. Fowler, B. L. Freeman: "The Los Alamos-Arzamas-16 High Magnetic Field Shot Series, Ancho Canyon Site, 1993", LANL report: LA-UR-94-2802. 2800 T
2003: C. M. Fowler, L. L. Altgilbers, Electromagnetic Phenomena (Ukraine) – review

2012: A. L. Velikovich – APS DPP tutorial talk, plus Phys. Plasmas 2015



# **Explosive flux compression: advanced configurations have been developed**



#### Strip generator powered railgun Ancho Canyon firing site (LANL)



Fig. 30. Strip generator powered railgun at Ancho Canyon firing site (LANL).



Fig. 31. FCG Driven railgun scheme proposed by Shvetsov et al. [132].



### Flux compression with lasers: Omega experiments with MIFEDS

- Cylindrical<sup>1,2</sup> and spherical<sup>3,4</sup> implosions
- CH capsule, D2 gas fill
- <= 8 T seed field
- T<sub>ion</sub> up by 15%, fusion yield up by 30% (spherical)



1 O. V. Gotchev et al., Phys. Rev. Lett. (2009); 2 J. P. Knauer et al., Phys. Plasmas (2010) 3 P. Y. Chang et al., Phys. Rev. Lett. (2011); 4 M. Hohenberger et al., Phys. Plasmas (2012)



## Fusion: magnetic, inertial, magneto-inertial





## Magnetic fusion

- T<sub>ion</sub> ~ 10 keV for DT, like ICF
- Need B to confine 3.5 MeV alpha:
  - B = 10 T  $\rightarrow$  R<sub>a</sub> = 2.7 cm
  - Plus exotic orbits, e.g. bananas
  - Machine size matters!
  - $\rightarrow$  Meter scale, > \$1 Billion cost
- Low beta: magnetic >> plasma pressure
  - beta < 0.1 for stability
  - B = 10 T, T = 10 keV  $\rightarrow \rho$  < 4E-10 g cm<sup>-3</sup>
- Steady state: Alpha heating balances losses
  - $P_{\alpha} + P_{aux} > P_{brem} + P_{dif}$
- Energy confinement time  $au_E$ 
  - $P_{dif} = n_e T_e / \tau_E$
  - Very complicated turbulent transport
  - Increases with B field!
- Lawson condition:  $n_{ion} T > f(T)$





## **Magnetic and inertial fusion scales**

Parameter	Magnetic	Inertial (hotspot)
N <sub>ion</sub> [cm <sup>-3</sup> ]	1E14	2E25
$\rho  [\text{g cm}^{-3}]$	4E-10	100
T <sub>ion</sub> [keV]	10	10
Pressure	1.6 bar	320 Gbar
Energy confinement time $ au_E$	10 sec	100 ps
System size	1 m	50 um
"Confinement length"	1E9 cm	50 um
$R_E = c_{sound} \tau_E$		
$ ho R_E $ [g cm <sup>-2</sup> ]	0.4	0.5
$ ho R_E T_{ion}$ [g cm <sup>-2</sup> keV]	4	5

"Confinement length" R<sub>E</sub> for MFE

- Convenient scale to compare with ICF
- >> system size as well it should be! B field is doing something



#### Lawson condition for magnetic and inertial fusion



#### "Triple product" pRT for self-heating

$$\begin{split} P_{\alpha} + P_{aux} + PpdV > P_{brem} + P_{dif} \\ P_{a} = f_{a} n_{a} E_{a} = alpha \ deposition \\ \text{MCF:} \ f_{a} = 1; \ pdV = 0 \ (\text{pressure balance}) \\ \text{ICF:} \ f_{a}(\rho R, T) < 1 \end{split}$$



#### Magneto-inertial fusion has a long history

#### Linhart et al., 1962

The principle is described of inertial confinement of plasma in which a cylindrical metallic shell compresses a magnetic field and plasma in a "soft-core" geometry. The shell transforms a part of its kinetic energy into the thermal energy of the compressed plasma and into the stored magnetic energy... J. G. Linhart, H. Knoepfel and C. Gourlan, "Amplification of magnetic fields and heating of plasma by a collapsing metallic shell," Nucl. Fusion Suppl. Pt. 2, 703 (1962).



MIRAPI – Minimum Radius Pinch

 Outer shell either ionized dust (MIRAPI-1) or tungsten 200-wire array (MIRAPI-2)

Inner shell hydrogen
 Still relevant for Z-pinch neutron
 production

#### Looks like MagLIF!

Slide courtesy A. Velikovich



## Magnetized ICF: help ignition, hurt burn?

"Magnetized ICF:" intermediate B field

- B field reduces electron conduction and alpha losses
- But magnetic pressure << plasma pressure</li>

Lindemuth and Kirkpatrick, Nucl. Fusion 1983:

- OD model, azimuthal B field Jones and Mead, Nucl. Fusion 1986:
- 1D model, azimuthal B field
- B field can help get to ignition
- BUT can impede propagating burn, high gain

#### From Jones and Mead:

Let us ask which properties would make a desirable regime in which to consider fusion with magnetized DT fuel. We would like the magnetic pressure to be negligible for a number of reasons: the magnetic field can impede the implosion, it can destroy the symmetry of the implosion, and in the  $\beta = 1$  regime the usual deleterious magnetic fusion energy (MFE) instabilities can occur. Here,  $\beta$  is the ratio of plasma pressure to magnetic pressure. One might like the conductivity and alpha mobility to be small in order to reduce losses to a pusher.





# Magnetic field reduces electron heat conduction, increases alpha deposition

#### **Electrons:**

- Axial field, spherical geometry
- Magnetized in 2 directions
- Radial heat flux 1/3 unmagnetized value

$$\langle q_r \rangle (B) = \frac{1}{3} \langle q_r \rangle (B = 0)$$
  
 $\langle q_r \rangle (B = 0) = -\kappa_{||}^e \frac{\partial T_e}{\partial r}$ 

#### Alphas: energy deposition in hotspot<sup>1</sup>



 $\alpha$  range:  $\rho R_{\alpha}[g/cm^2] \approx 0.44 \ (T_e/10 \text{ keV})^{5/4}$  Strong B:  $\alpha$  cyclotron freq > rate  $\alpha$ 's stop on e-'s

$$B[T] = 7850 \frac{\rho [100 \text{ g/cm}^3]}{T_e [10 \text{ keV}]^{3/2}}$$

Example:

T<sub>e</sub> = 10 keV,  $\rho$  = 100 g/cm<sup>3</sup> → B = 7850 T, or B<sub>0</sub> = 8.7 T at conv. ratio 30



# Imposed magnetic field expands hotspot self-heating domain





## Imposed field improves capsule performance

"Perkins approach:" recovery of 1D ignition

- Target ignites in 1D
- But not in lab
- Perturb simulation to degrade yield
- Imposed B: ignites for stronger perturbation
- Does not increase 1D yield

#### Too much of a good thing: $B_0 = 100 T \rightarrow pancaked implosion$



#### See also:

Perkins et al., Phys. Plasmas 2017 LDRD final report – contact Perkins

Perkins and Ho support both approaches, just a mnemonic



"Low foot" NIF design: imposed axial B



2D RZ Lasnex MHD simulations L. J. Perkins et al., Phys. Plasmas 2013



### MagLIF<sup>\*</sup> = Magnetized Liner Inertial Fusion: fusion gain on device like Sandia's Z Machine

#### MagLIF schematic S. A. Slutz, IFSA 2015 talk



\*S. A. Slutz, R. A. Vesey, Phys. Rev. Lett. 2012

#### 3 key elements A. B. Sefkow, Phys. Plasmas 2014





- Laser preheat: ~ 10 kJ, ions to ~ 500 eV
- $B_{z0} \simeq 30$  T lower e- thermal loss, confine alphas
  - Relaxed hotspot ρR
  - Too much field impedes burn propagation
- Slow implosion: liner v<sub>imp</sub> <~ 100 km/s</li>
- Long implosion time ~ 60 ns
  - Laser preheat over ~10 ns: low power, intensity



- Draw a self-heating curve for MagLIF's hotspot.
- Place it on the plot from Darwin Ho on slide 50.
- Recall that MagLIF is a cylindrical and not spherical system.



B fields in hohlraums

- Self-generated
- Imposed



## Self-generated fields in hohlraums<sup>1</sup>: Highly localized ~ MG fields, reduced heat flow

- Main effect on plasma: reduced electron heat conduction perpendicular to B
- NIF shot N151122: HDC, 0.3 mg/cc hohlraum fill late peak power [5 ns]



HYDRA MHD 2D R-Z simulations

- B field advection, diffusion, Biermann battery, Nernst effect, reduced heat flow
- No Righi-Leduc effect but now in code [J. Koning]

<sup>1</sup>W A Farmer, J M Koning, D J Strozzi, D E Hinkel, L F Berzak Hopkins, O S Jones, M D Rosen, Phys. Plasmas (2017)

Lasnex work by M. D. Rosen: similar results, but Nernst more fully "erases" Biermann fields



## Self-generated fields in hohlraums<sup>1</sup>: increased T<sub>e</sub>, later bangtime, less pancaked shape



<sup>1</sup>W A Farmer, J M Koning, D J Strozzi, D E Hinkel, L F Berzak Hopkins, O S Jones, M D Rosen, Phys. Plasmas (2017)



## Self-generated fields in hohlraums<sup>1</sup>: Nernst effect advects field into wall, reduces T<sub>e</sub>

High-foot simulation, Late peak power [13 ns] Electron temperature [keV]



• Nernst effect: B field advected by e<sup>-</sup> heat flow

- Reduced by things that limit heat flow:
  - Nonlocality, ion acoustic turbulence
- MHD sims with fluid Nernst: likely upper bound on effect

"What Biermann giveth, Nernst taketh away" – M. D. Rosen

<sup>1</sup>W A Farmer, J M Koning, D J Strozzi, D E Hinkel, L F Berzak Hopkins, O S Jones, M D Rosen, Phys. Plasmas (2017)





#### **Self-generated B fields in capsule**

Two recent works: effects are there but not huge:

- Walsh et al., Phys. Rev. Lett. 2017
- M. Partha (summer student), S. Haan LLNL 2016



## Imposed B field effects on hohlraum dynamics<sup>1</sup>

- Underdense hohlraum fill magnetized for  $B_{z0} \approx 10 \text{ T}$
- Raises plasma temperature
- Improves inner-beam propagation:
  - Higher T<sub>e</sub>, lower n<sub>e</sub> in equatorial channel → less absorption
- May reduce SRS<sup>2</sup>
- Hot electrons:
  - Magnetized in underdense fill
  - Could guide toward or away from capsule
  - Strongly depends on when and where sourced

<sup>1</sup>D J Strozzi, L J Perkins, M M Marinak, D J Larson, J M Koning, B G Logan, *J. Plasma Phys.* (2015)
 <sup>2</sup>D S Montgomery, B J Albright, D H Barnak, P Y Chang, J R Davies, et al., *Phys. Plasmas* (2015)





## Imposed B field raises plasma temperature, improves inner-beam propagation







# Inner beam propagation: B field reduces inner beam absorption in fill, less pancaked implosion



Limited Distribution

# B field follows MHD "frozen-in law:" advected with flow

Low-foot shot N120321 B<sub>z0</sub> = 70 T 18 ns: early peak power



Light blue boxes are hot electron sources for next slide

<sup>1</sup>D J Strozzi, L J Perkins, M M Marinak et al., J. Plasma Phys. (2015)



### Hot electrons: coupling to DT early in peak power is very sensitive to source location





## **Outline** again

#### **FUNDAMENTALS**

- Particle motion, drifts, magnetic mirroring ٠
- Waves in magnetized plasma, Faraday rotation ۰
- MHD single-fluid model: derivation ۰
- MHD effects •
  - Magnetic pressure
  - Reduced heat flow •
  - Frozen-in law
  - Biermann battery

**OTHER DOMAINS not covered** 

Astrophysics, lab-astro

Nernst, Righi-Leduc •

## Thank you!

#### **EXPERIMENTS / APPLICATIONS**

- Diagnostics
- Imposed fields pulsed power, laser-driven coils
- Flux compression explosives, laser driven
- Fusion •
  - Magnetic vs. inertial fusion
  - Magneto-inertial fusion
  - Imposed fields in ICF
  - MagLIF
- B fields in hohlraums: self-generated, imposed



Plasma propulsion



