

NIC



LPI experiments with single and multiple NIF beams

Anomalous Absorption Conference
June 2012

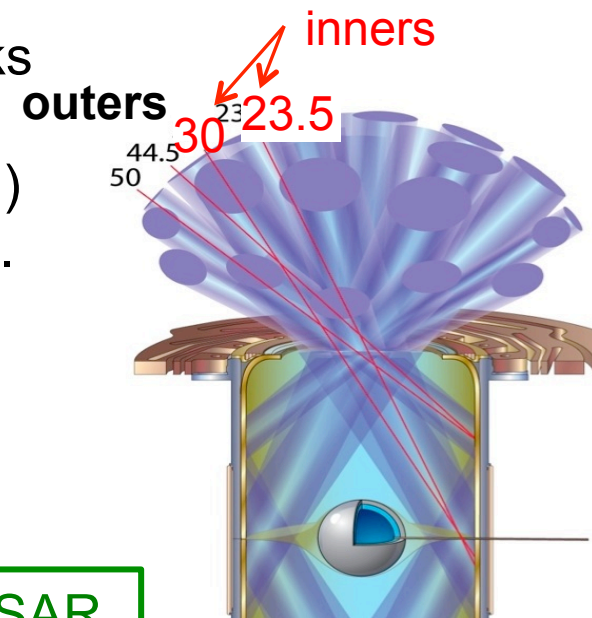
**D. J. Strozzi, J. D. Moody, H. F. Robey, L. Divol, P. Michel,
R. L. Berger, E. A. Williams, D. E. Hinkel**
Lawrence Livermore National Lab

Lawrence Livermore National Laboratory • National Ignition Campaign

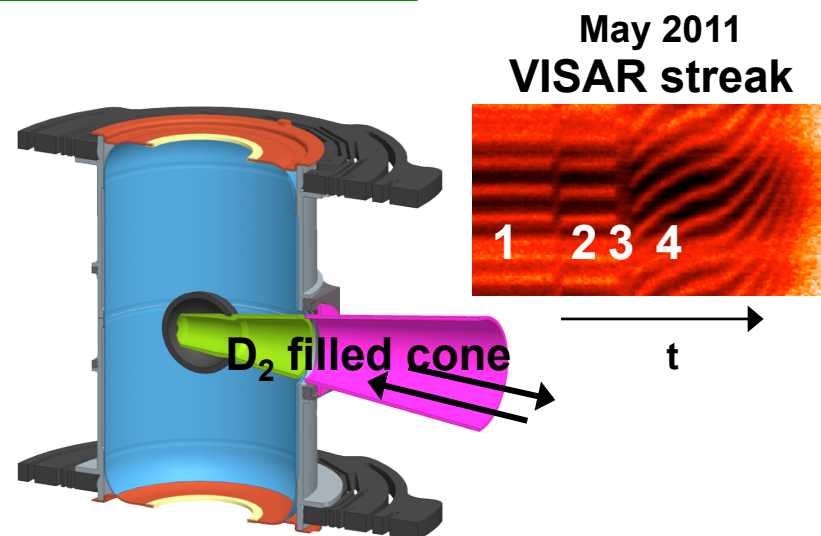
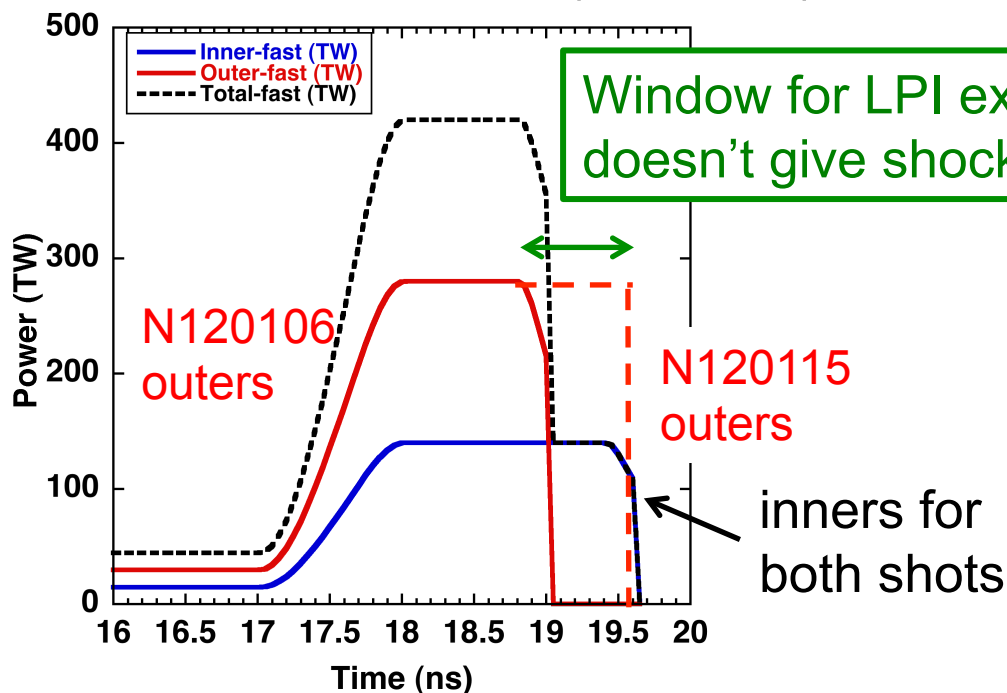
This work performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344

N120106 and N120115: “keyhole” shots isolate effect of outer cones on inners

- “Keyhole” targets: VISAR detector: tune first 3 shocks
- “Fast rise” (~1 ns) from third to fourth shocks
- Gold hohlraum with “large” laser entrance hole (LEH)
- 3 laser colors: $\lambda_{30}-\lambda_{out} = 6.6$ Ang., $\lambda_{23}-\lambda_{30} = 1.5$ Ang.
- Same energy, except outers extended in N120115
- Very low inner-cone SBS

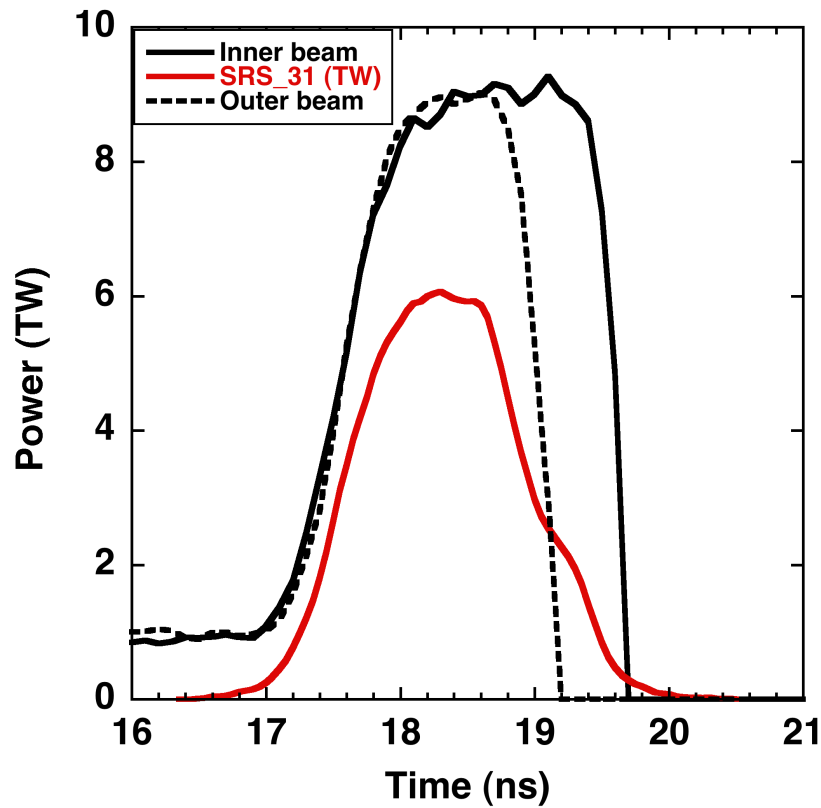


Fourth pulse fast-rise (N120106)

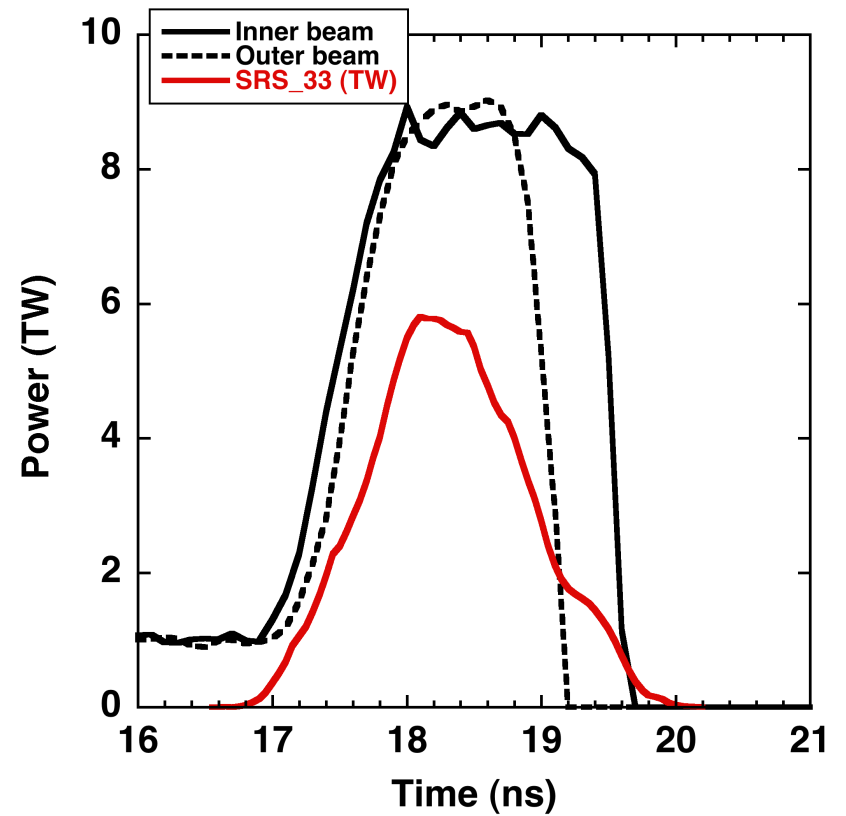


N120106: SRS on both inner cones drops when outers turn off

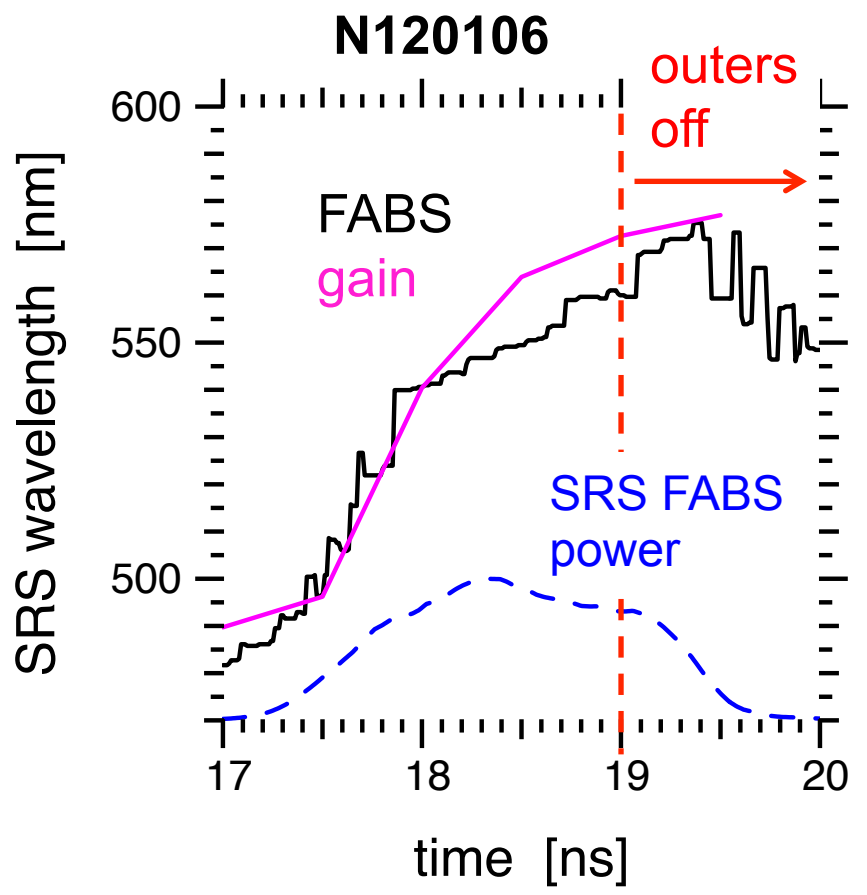
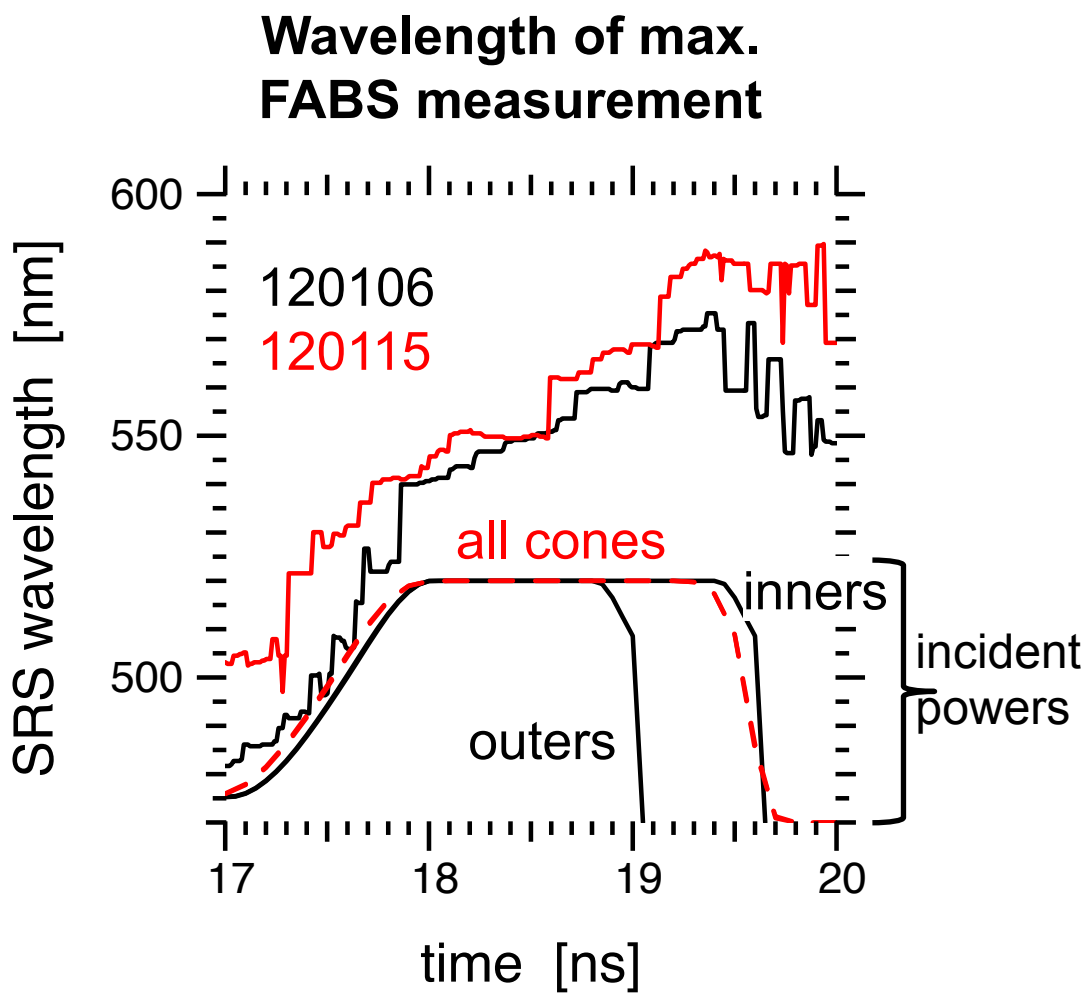
30° quad



23° quad



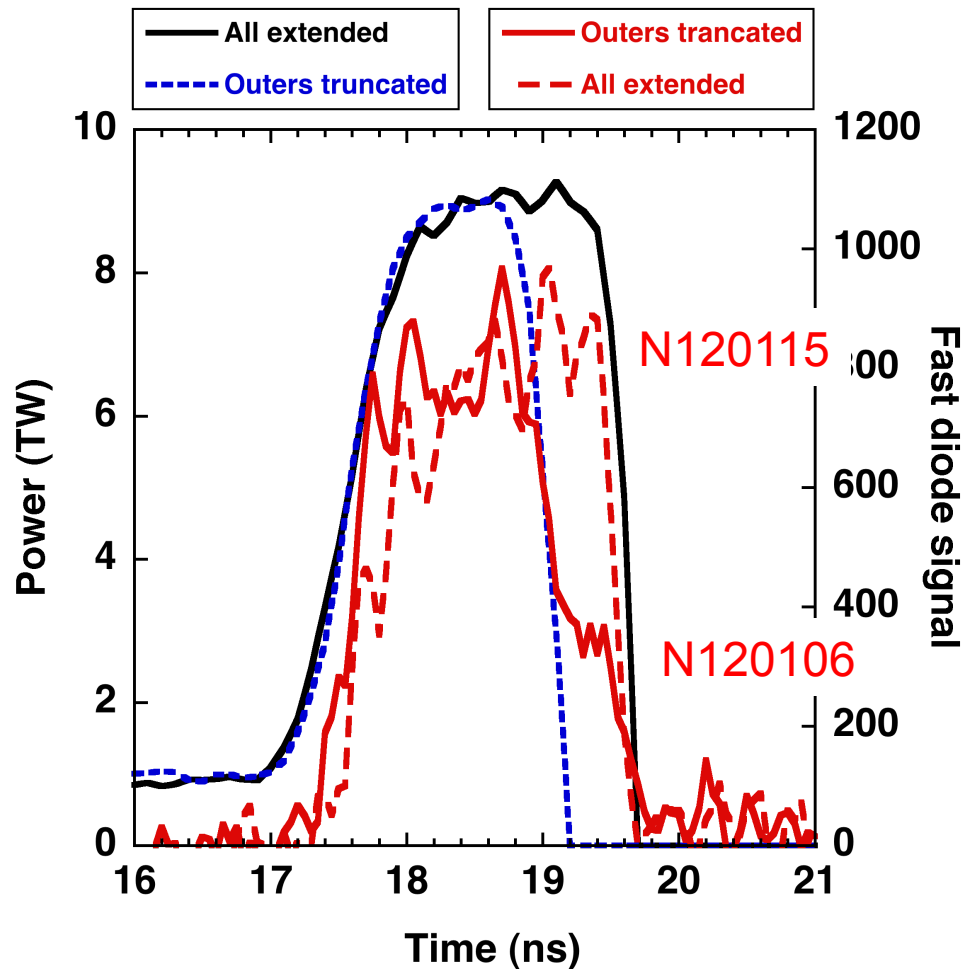
SRS spectrum on 30° cone evolves similarly if outers truncated or not; gain spectrum within ~10 nm



Gain spectrum from Hydra rad-hydro by H. Robey with high-flux model; low-flux model gives 20-30 nm longer wavelength

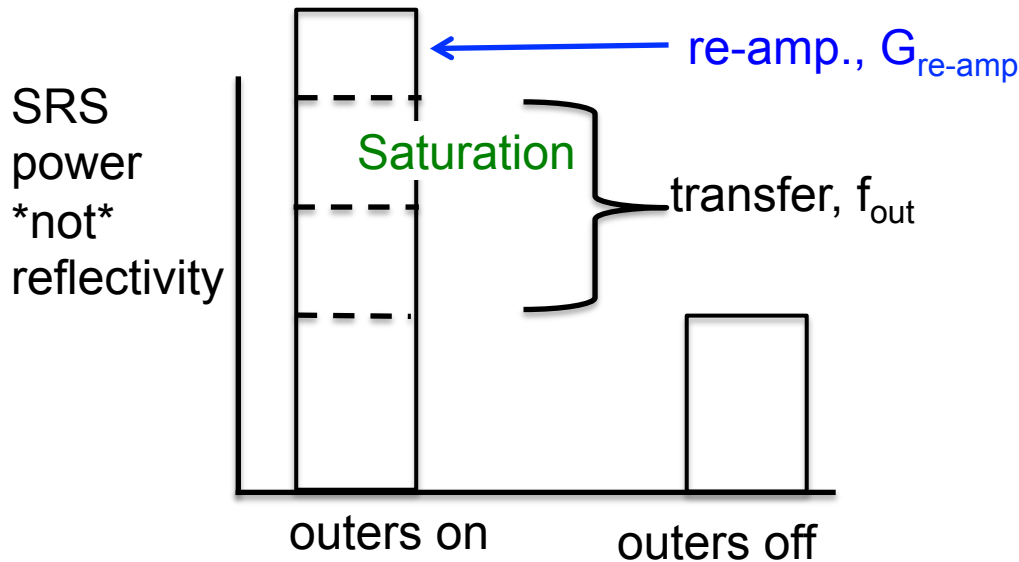
N1201{06, 15}: SRS power when outers are off is 40% of value with outers on

SRS data for 30° cone

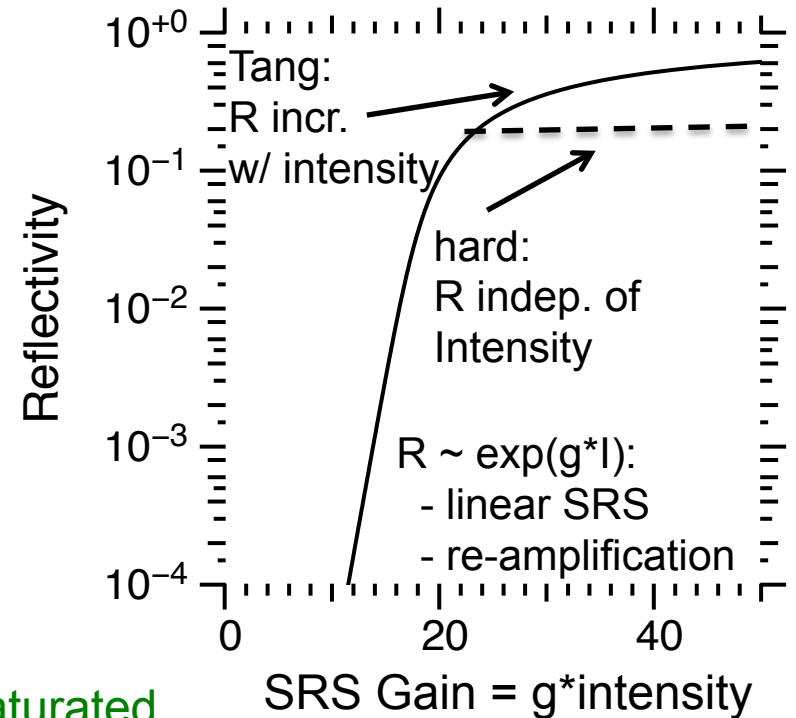


30° cone SRS power: N120106 (outers off) / N120115 (outers on): **0.4**

Inner SRS drops when outers off: no power transfer, no re-amplification, less saturation



Linear gain; Tang vs. hard saturation



When outers off:

Decreases because transfer stops

Decreases if Tang (not hard) saturated

Decreases because re-amplification stops

$$P_{30}^{SRS} = P_{30}^{inc} \cdot R(g * I_{30}) \cdot R_{re-amp}$$

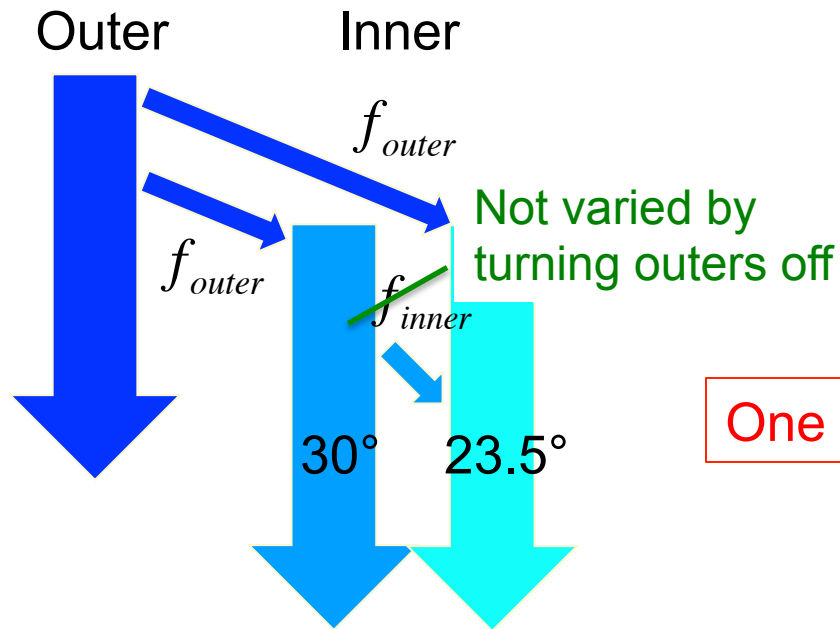
$$P_{30} = P_{30}^{inc} + \frac{1}{2} f_{out} P_{out}^{inc}$$

$$R_{re-amp} = \exp[G_{re-amp}]$$

Two unknowns: f_{out} , G_{re-amp}

SRS light from one beam amplified when crossing another beam [P. Michel et al.]

One-parameter model for cross-beam energy transfer from outer to inner cones



One unknown

$$P_{23,30} = P_{23,30}^{inc} + \frac{1}{2} \Delta P_{out}$$

ΔP_{out} = total power xfer from outer to inner cones

$$= f_{out} P_{out}^{inc}$$

$$P_{out}^{inc} = P_{44}^{inc} + P_{50}^{inc} \approx 4 P_{30}^{inc}$$

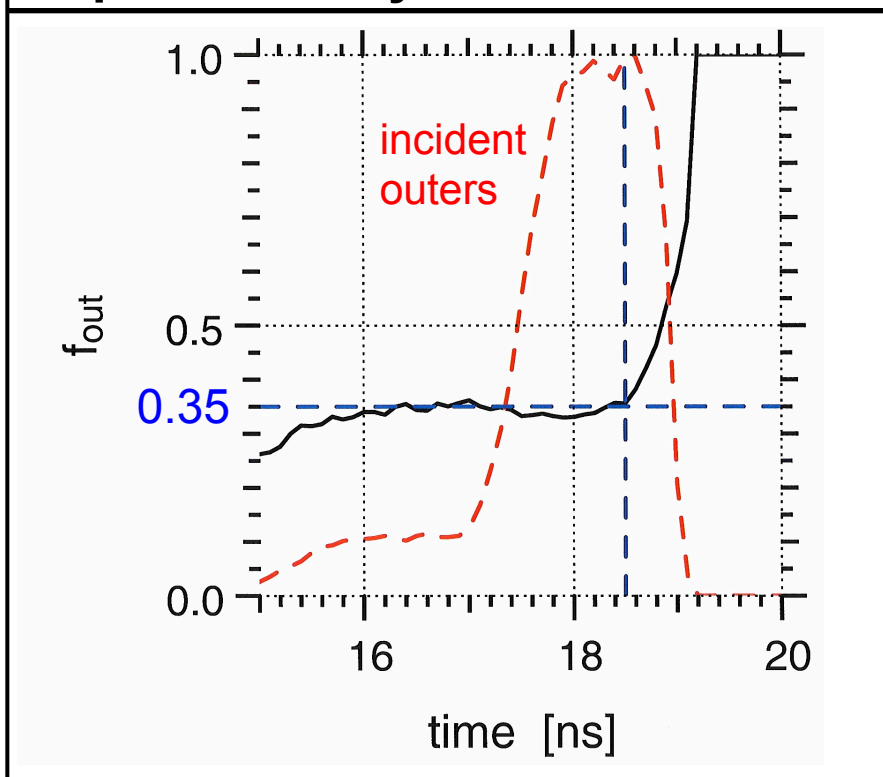
$$P_{23,30} = P_{23,30}^{inc} (1 + 2 f_{out})$$

f_{out} = fraction of outer-cone power transferred to inner cones

We neglect spatial non-uniformity in transfer, which should be looked at (R. L. Berger)

Measured ratio of SRS powers relates transfer and re-amplification

Power transfer from H. Robey's post-shot Hydra of N120106

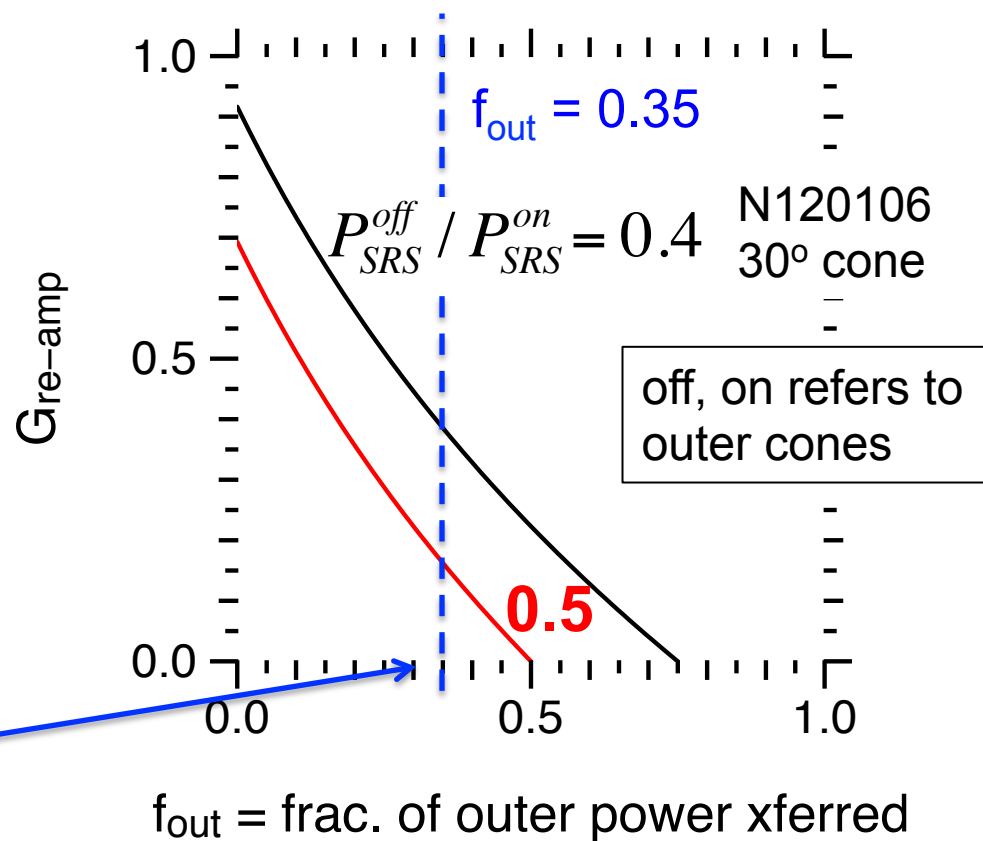


- $f_{out} \sim 0.35$ agrees w/ capsule symmetry data (R. Town)
- Hydra's f_{out} passes basic check: inconsistent w/ SRS drop if $f_{out} > G_{re-amp} = 0$ values

$$\frac{P_{SRS}^{on}}{P_{SRS}^{off}} = \rho \cdot (1 + 2f_{out}) \exp G_{re-amp}$$

$$\rho = \frac{R(g * I_{30}^{on})}{R(g * I_{30}^{off})} \rightarrow 1$$

hard saturation:
maximizes
 G_{re-amp}



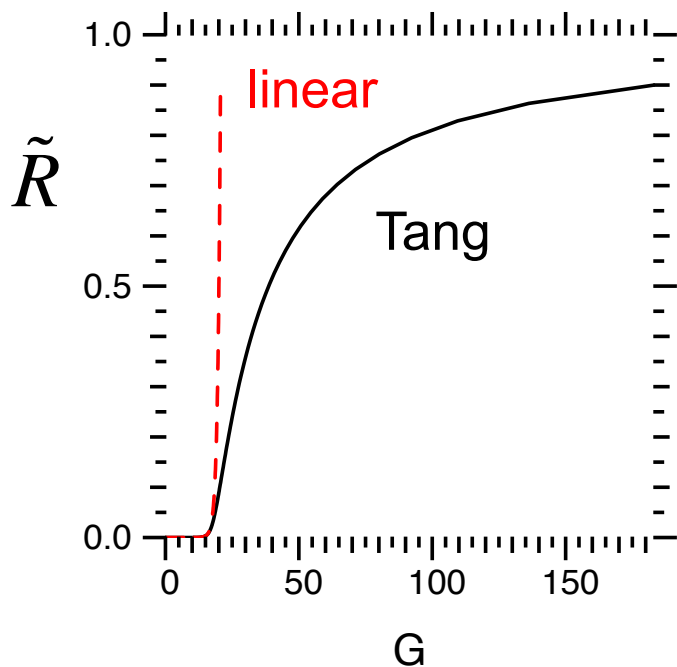
Allow for saturation to vary with intensity

- $G_{\text{re-amp}}=0$: SRS changes only due to power transfer; $G_{\text{re-amp}}>0$ lowers A, raises G
- Neglect change in g = role of plasma conditions in gain
 - measured spectra similar for N120106 and N120115

$$A \equiv e^{-G_{\text{re-amp}}} \frac{P_{\text{SRS}}^{\text{on}}}{P_{\text{SRS}}^{\text{off}}} = \rho(1 + 2f_{\text{out}}) = \alpha(f_{\text{out}}, G)$$

$$\rho = \frac{R(P_{30}^{\text{on}})}{R(P_{30}^{\text{off}})} = \frac{R(G \cdot (1 + 2f_{\text{out}}))}{R(G)}$$

$$G \equiv gI_{30}^{\text{inc}}$$



**Tang formula:
coupled-mode eqs.
w/ pump depletion**

$$\tilde{R}(1 - \tilde{R} + \tilde{s}) = \tilde{s} \exp[G(1 - \tilde{R})]$$

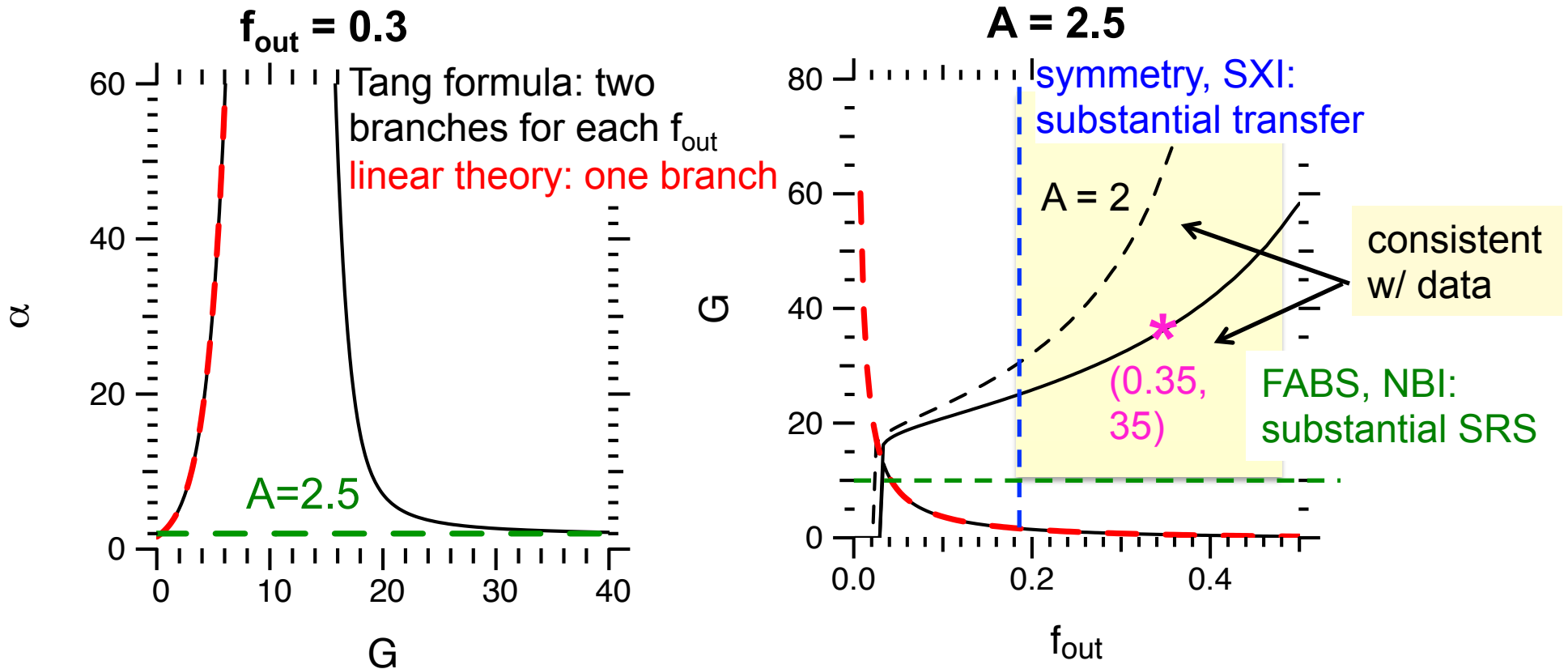
$$\tilde{R} = \frac{\omega_0}{\omega_1} R \quad \tilde{s} = \frac{\omega_0}{\omega_1} \frac{I_{1\text{seed}}}{I_0} \rightarrow 10^{-9}$$

linear theory

$$R = s \exp G; \quad G = \frac{1}{2f} \ln \left[\frac{A}{1 + 2f} \right]$$

Manley-Rowe

Inner-cone SRS is strongly saturated



- Small drop in SRS for large transfer:
 - Linear branch requires small gain, not consistent w/ large SRS
 - Consistent w/ large gain on strongly-saturated, pump-depleted Tang branch
- Estimating G and f_{out} :
 - N120106: SRS power $\sim 28\%$ incident when outers off; wavelength = 570 nm
 - No re-absorption of scattered light: $R_{\tilde{t}} = 0.45$, $G = 35 \rightarrow f_{out} = 0.33$ – near sims!
 - $R_{off} / R_{on} = 0.45 / 0.67 = 0.67$; $I \cdot dR/dI = 0.2$

Conclusions for N120106 and N120115: effect of outer cones on inners

- Inner cone SRS power approximately **doubled** by presence of outer cones
- Hydra modeling, and symmetry scaling with $\Delta\lambda$, suggest 35% of outer beam power transferred at time outers shut off
- Measured SRS decrease consistent with this 35% transfer, and with modest outer-inner re-amplification gain exponent of at most 0.15 - 0.4
- Saturation: neglecting re-amp. (which *minimizes* saturated gain): {large SRS, large transfer, and small SRS drop when outers} imply SRS is strongly saturated, not in linear regime

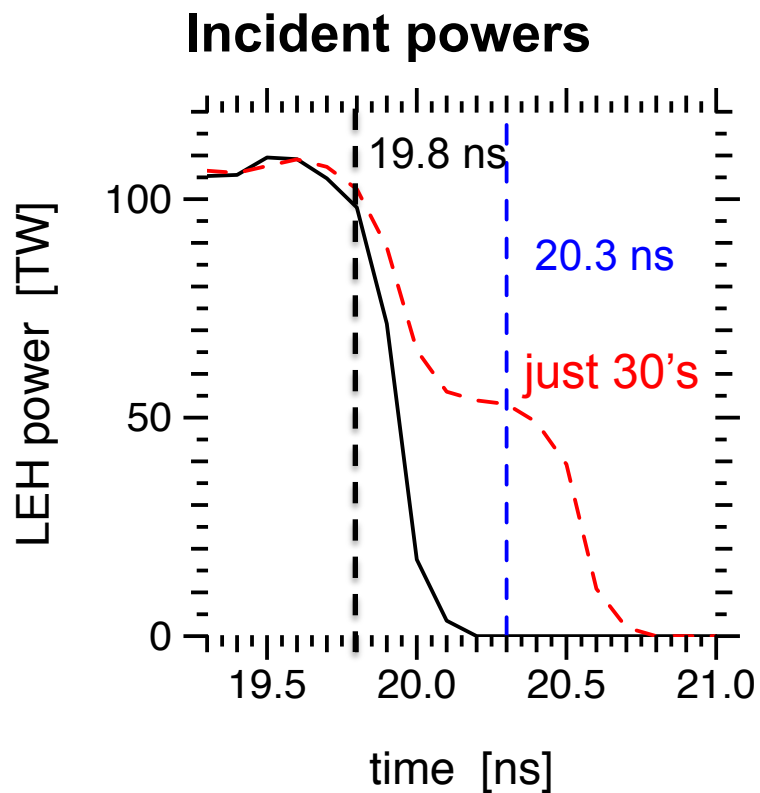
March keyholes: one inner cone extended – “single-quad” LPI experiments!

- But in a rapidly evolving, maybe azimuthally-asymmetric, hohlraum...
- Slow (3 ns) rise of fourth pulse
- “small” LEH
- 3-color scheme: $\lambda_{30} - \lambda_{\text{out}} = 7.3 \text{ Ang.}$; $\lambda_{23} - \lambda_{30} = 1.2 \text{ Ang.}$
- Single-quad expt. if no re-amplification of one inner quad by others on same cone
 - Power transfer excluded by azimuthal symmetry

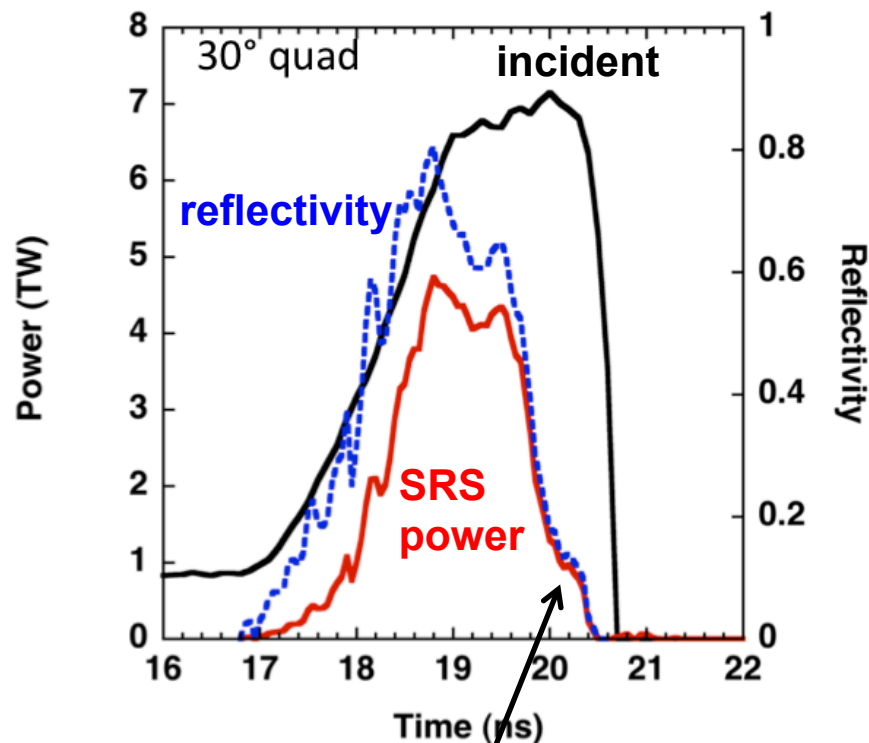
shot	hohlraum	peak power [TW]	extended cone
N120229	DU	420	30
N120303	DU	420	23
N120304	Au	420	30
N120305	DU	320	30

← shot we study

N120305: “single-quad” 30° cone SRS ~ 15-20%



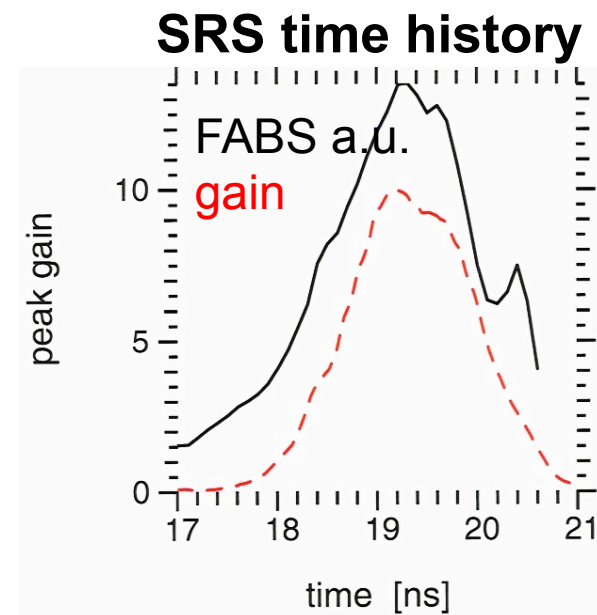
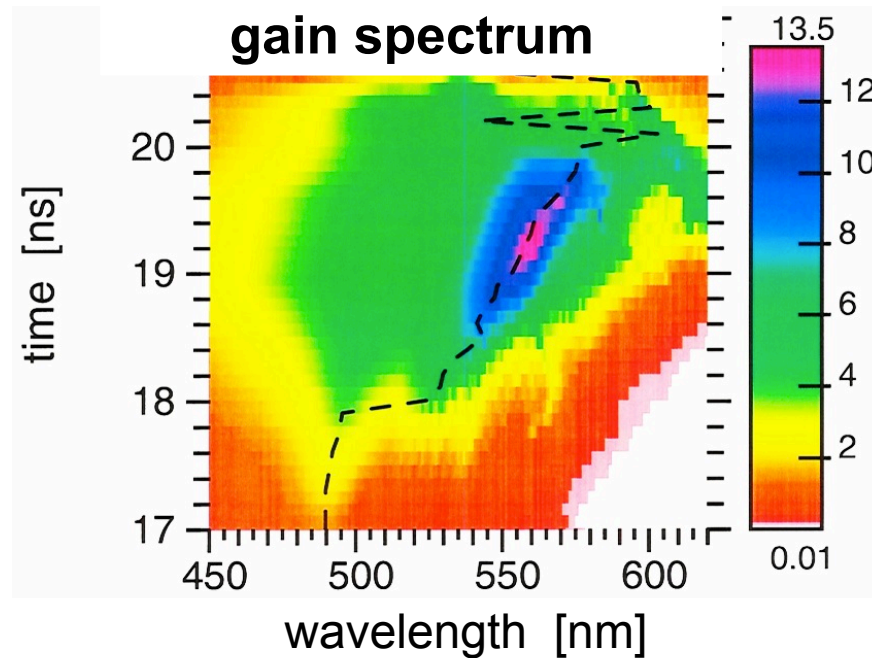
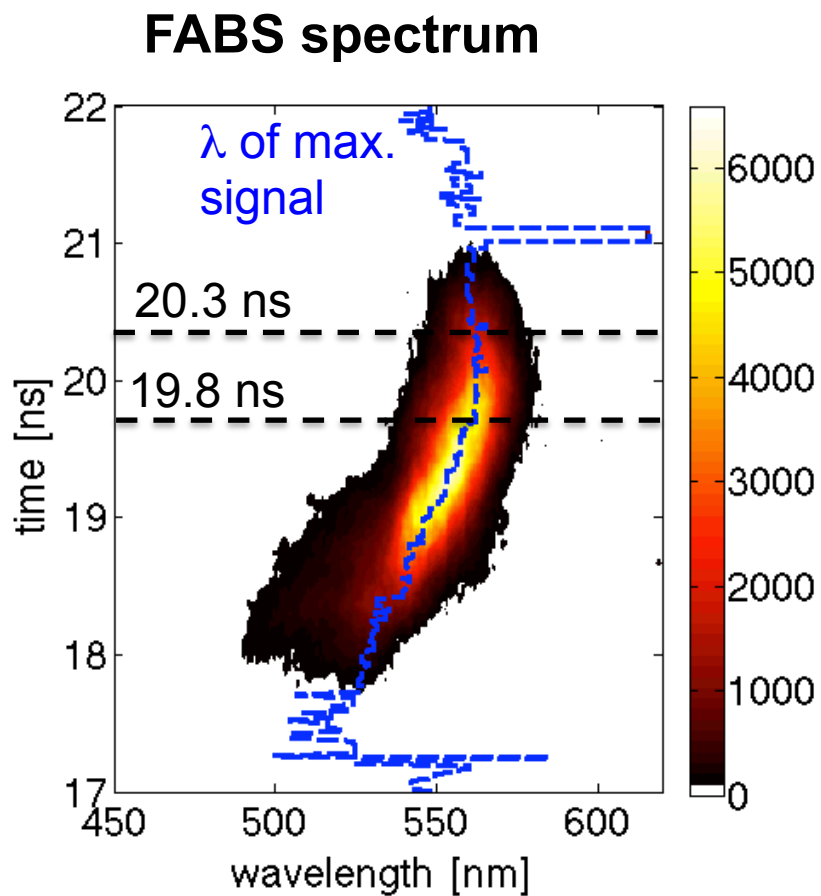
SRS from deconvolved NBI movie



plateau when just 30's on; ~15-20% SRS reflectivity depending on shot

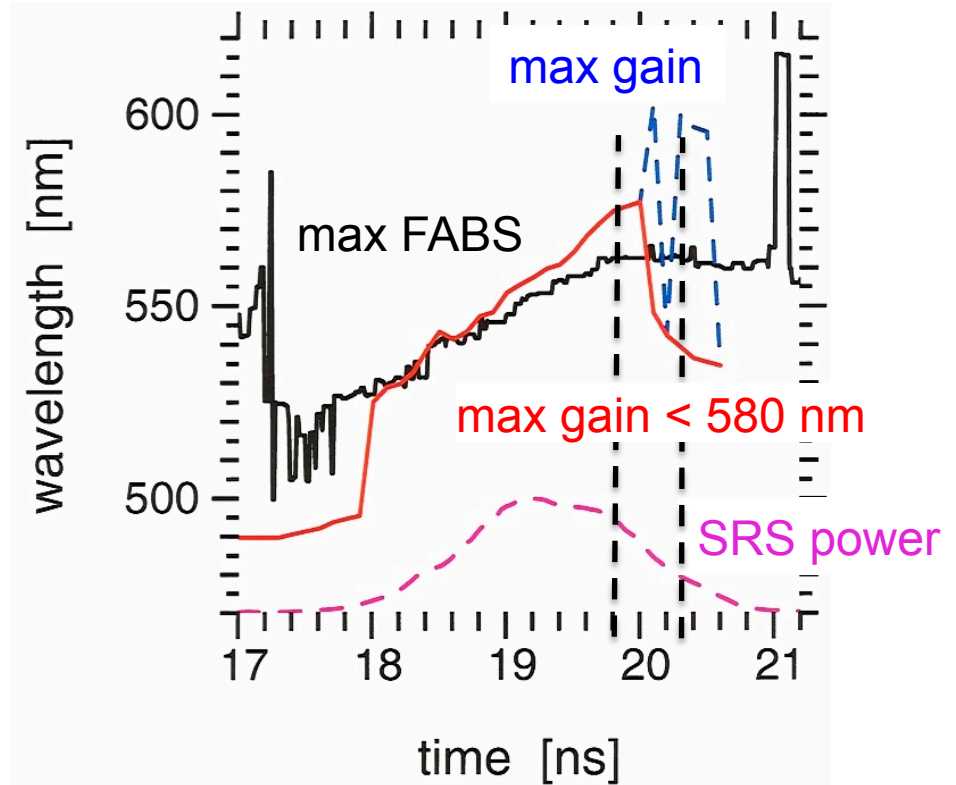
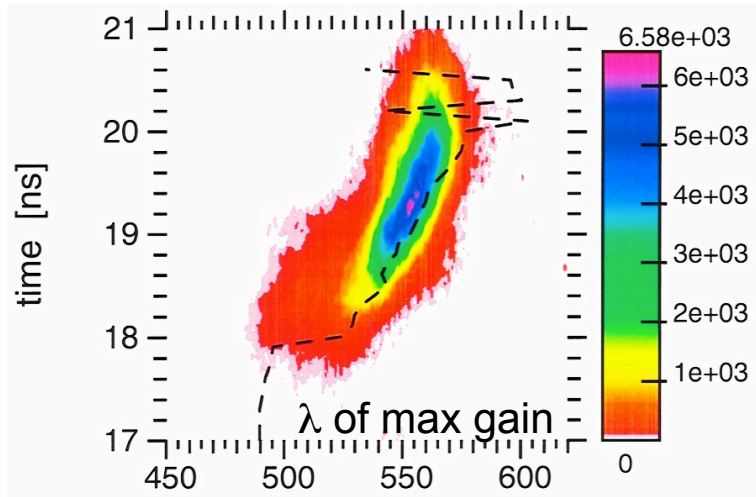
SRS power starts dropping before other cones turned off – complicates analysis

N120305: SRS spectrum on 30° cone

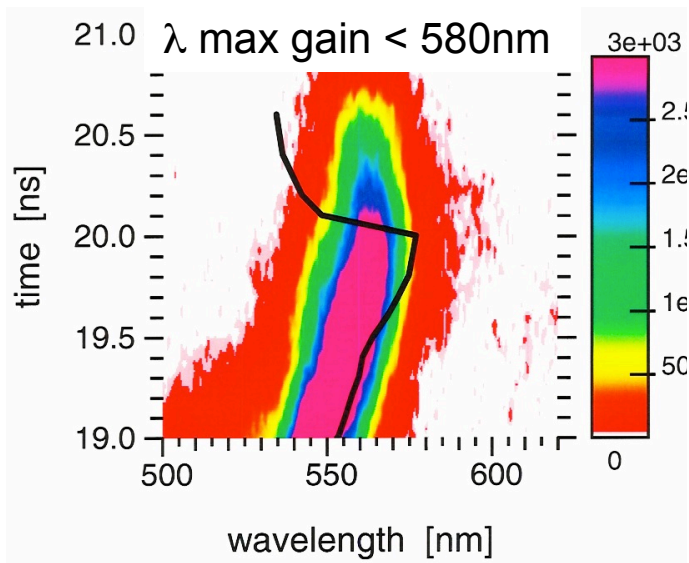


N120305: 30° cone SRS: FABS and linear gain agree when all cones on, differ when just cone 30 on

FABS spectrum

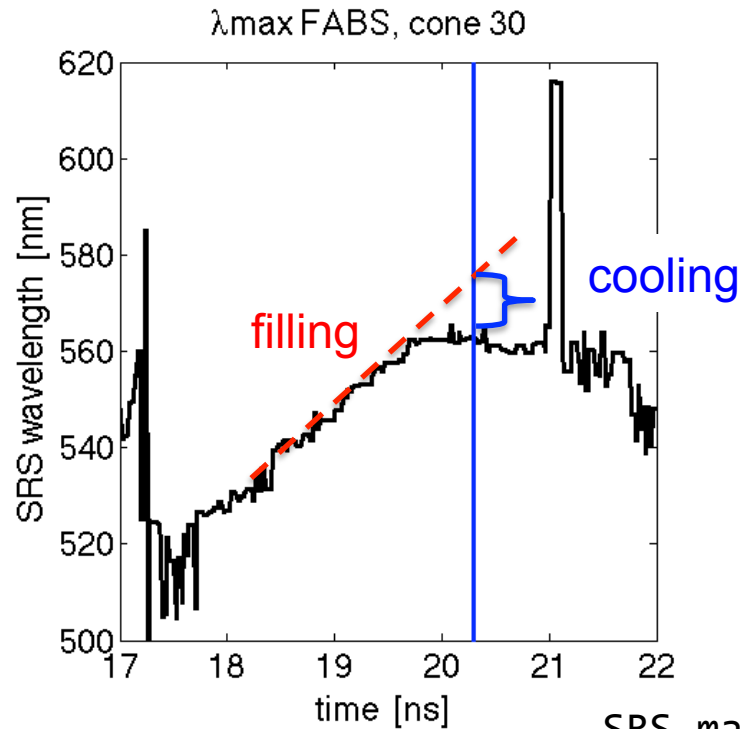


zoomed late, and capped



Multi-quad effects not essential to match spectrum

N120305: 30° cone SRS FABS spectrum: filling and cooling rates



SRS matching conditions: $\lambda, T_e \rightarrow n_e$

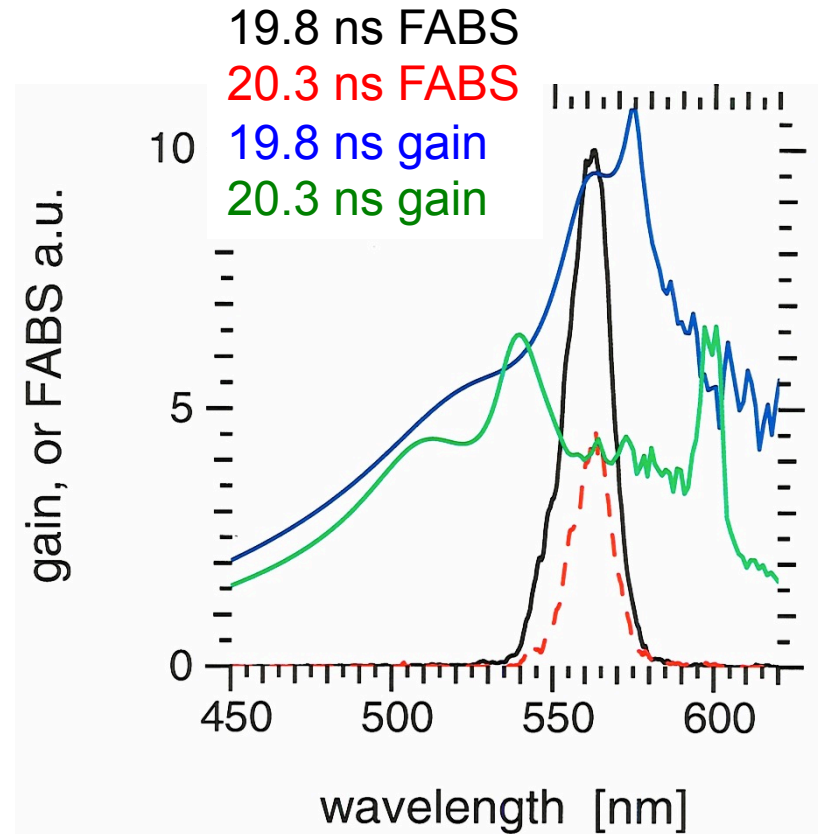
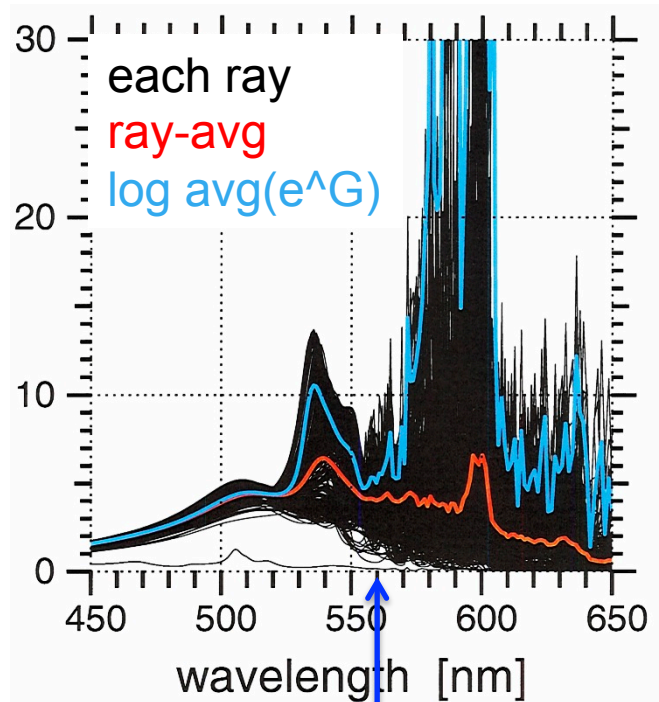
t [ns]	λ [nm]	n_e/n_{cr} [1.5 keV]	n_e/n_{cr} [3 keV]
19	545	0.102	0.069
19.5	558	0.115	0.082

fill rate: $dn_e/dt = 0.026 n_{cr}/ns$ for 1.5 or 3 keV!

cooling rate: $0.1n_{cr}$ and 2.3 keV match with 560 nm.
 To keep wavelength constant, $dn_e/dt = 0.026$ is
 balanced by $dT_e/dt = -1.1$ keV/ns

N120305: 30° cone SRS: what happens when other cones turn off

SRS gain, $t = 20.3$ ns



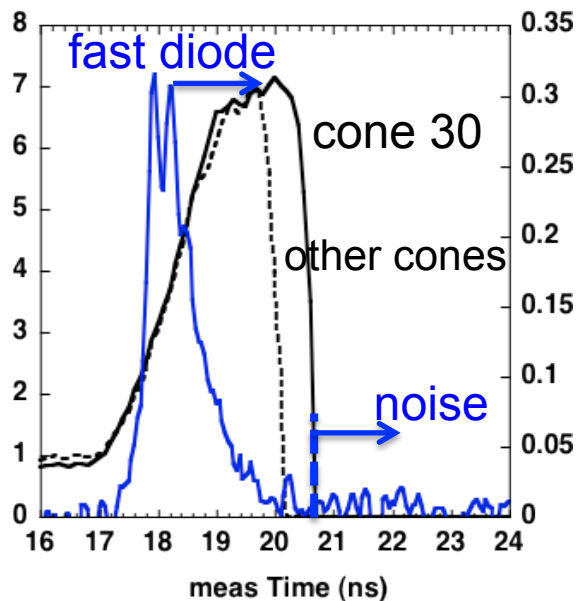
- FABS wavelength of 560 nm in “dead zone” of gain spectrum:
 - LPI physics (speckles, inflation) not likely to fix that
 - Plasma conditions likely wrong

Why gains blueshifted at late time?

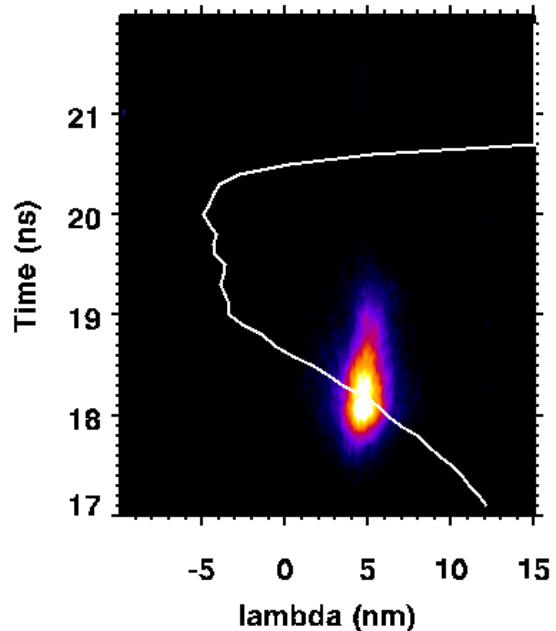
- Azimuthal non-uniformity:
 - Plasma really hotter in beams?
- Rad-hydro codes have trouble when laser turned off?

N120305: SBS on 30° cone: gains ~ 1.5 Ang. redshifted vs. measurement

SBS power [TW]

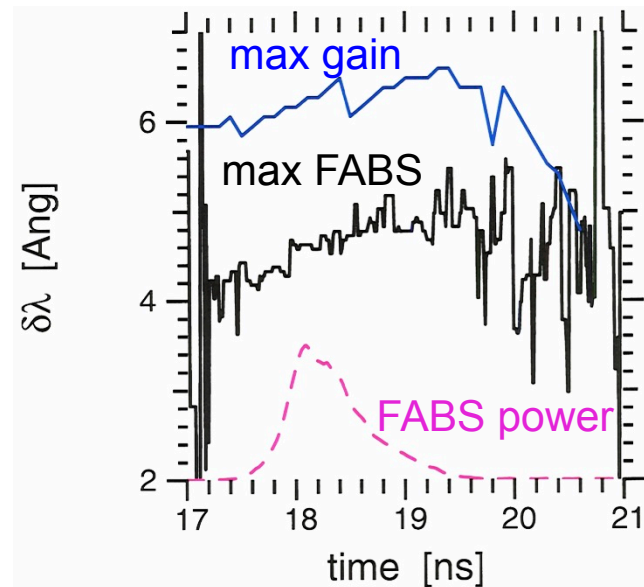
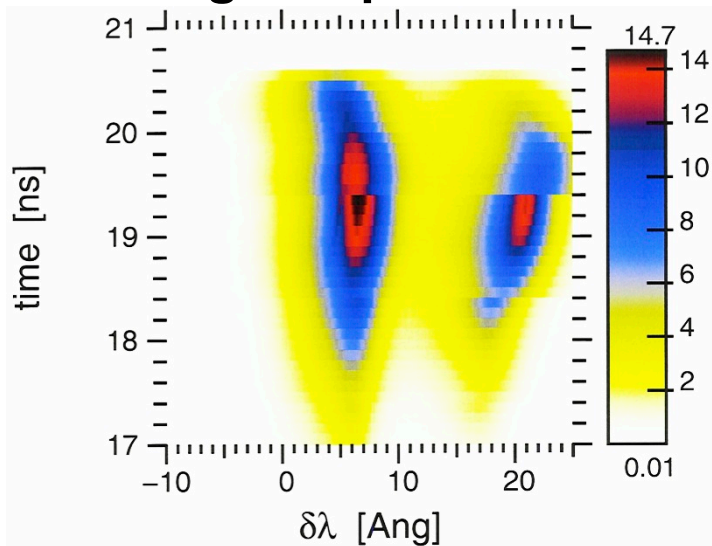


FABS SBS spectrum



Time-dependent n_e^1 may explain redshift [R. L. Berger]
¹T. Dewandre, J. R. Albritton, E. A. Williams, Phys. Fluids 1981

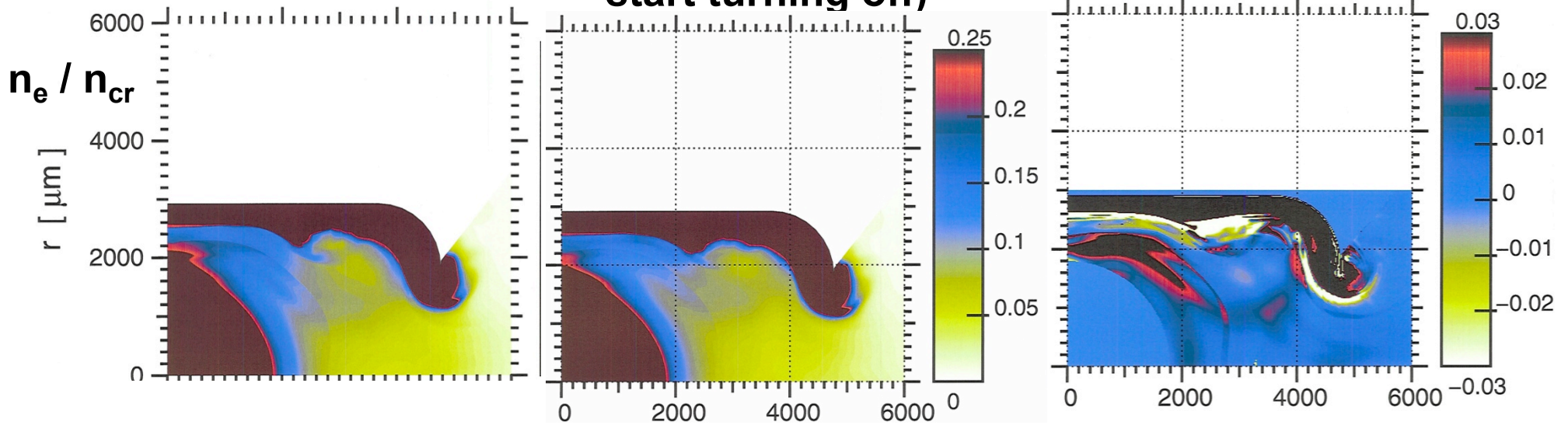
SBS gain spectrum



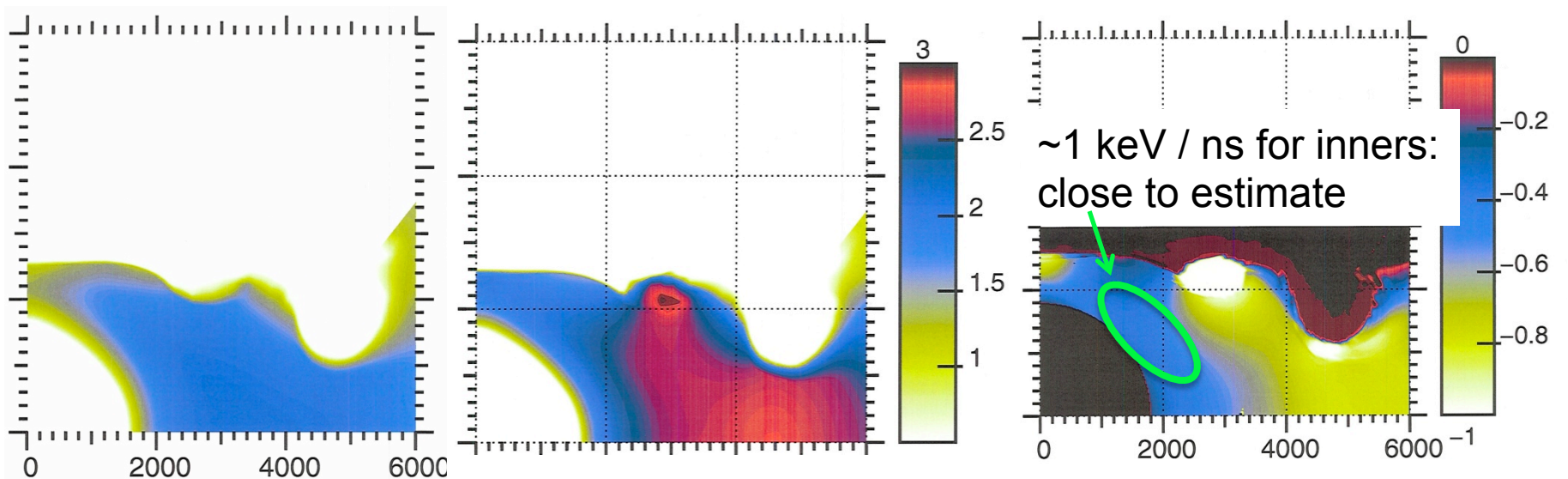
N120305: plasma maps from H. Robey's R-Z post-shot Hydra runs

20.3 ns (just 30's on) - 19.8 ns (other cones start turning off)

= difference



T_e [keV]



Conclusions on March 2012 keyholes

- Measured single-quad inner-cone SRS ~ 15-20%
- When all cones on, single-quad gain spectrum close to data
 - multi-quad effects not essential
- Gain spectrum changes more when just 30's are on than data –
 - rad-hydro plasma conditions likely wrong, perhaps due to degraded drive
- SBS gain spectrum ~ 1.5 Ang. redshifted vs. data
 - time-dependent density may explain